# MATHBMATIGAI TREATMENT 

OF

## DATA CONTAINTNG

INSTRUMCNI RESPONSE

# MATHBMATICAL TREATMENI OF DIGITIZED DATA CONTAINING INSTRUMENT RESPONSE AND STATISTICAL DEVIATIONS 

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SCOPE AND CONTIENTS:
Two methods of approack are presented for the analysis of digitized data containing instrument response effects. The first method corrects the data vector by means of multiplication by the inverse response matrix. In order to accomodate large size data fields, special effort is made to obtain expressions for the inverse matrix elements in closed form. Reducto ion of statistical uncertainties is accomplished by application of nomnegativity conditions. The second approach is based on the method of least squares. Applications to two-dimensional coincidence spectra and nonlinear model functions are discussed in some detail. Although the main ewphasis is placed on analysis of nuclear apectra, the techniques presented need not be limited to this application alone.

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## CHAPTER I

INIRODUCTION

Becent developments in nuclear spectroscopy bave flourished to a high degree of sophistication in data acquisition equipment and techniques. Multi-channel analyzers with up to 16,000 channels are becoming increasingly common. As a consequence of the sheer bulk of data generated the experimenter must perforce have a less intimate contact with raw results and take recourse to outomatic data processing. Fortunately this approach is made possible by the increasing availability of high speed and large memory capacity computers.

The physically meaningiul parameters are in general not directly arailable from the raw data. Usually it is necessary to create a Eamily of mathemacical models for the spectrum and thea apply a process of selecting a particular model which will adequately describe the observed data. The required physical quantities are then contained within the parameters of this model. Construction and mathematical methods of prom cessing such models will be the main subject of this thesis.

The nuclear spectra considered bere possess two particular characteristics which rexder theis analysis difficult. Firstly, the detector response to mono-mergetic radiation is usualiy a compicated sunction which may contain more than one peak. Secondly, the number of counts accumulaced in each analyzer channel is subject to statistical deviations. As a consequence, the observed spectrum way bear little
similarity to the actual distribution of energies emanating from the nuclide under study. An example of such response is given in Figure 1 which shows the ${ }^{137} \mathrm{Cs}$ r-ray spectrum taken with a $\operatorname{NaI}(T \ell)$ scintillator. This particular muclide decays by $\beta^{-}$emission to ${ }^{137}$ Ba which subsequently deexcites by emitting a $662 \mathrm{keV} \gamma$-ray. The main peak shown in channel 157 is produced by the photo-electric effect which usually results in deposition of all incident $\gamma$-ray energy inside the scintillator. Structure to the left of this peak is due to partial energy losses. The plateau below channel 110 is caused by Compton scattering events after which energydegraded photons may escape from the detector. Compton-scattered $\gamma$-rays from material outside the scintillator produce the small peak at channel 46. The entire response displays effects of a Gaussian resolution which is caused in part by statistical fluctuations of the number of electrons collected in the photomultiplier. This effect presents the greatest difficulty when the spectrum contains $\gamma$-rays with small energy separation. Most types of response functions considered in this thesis will take the Gaussian resolution into consideration.

It is possible to derive a general mathematical expression for the observed spectrum. Let the detector response function to incident radiation of energy $y$ be given by $R(x, y)$. If we stipulate the condition of normalization

$$
\int_{0}^{\infty} R(x, y) d x=1
$$

then we can think of $R(x, y) d x$ as the probability that an event of energy $y$ in the true spectrum will produce a response in the energy interval $x$ to

to $x+d x$. The concept of the "true" spectrum $T(y)$ as used in this context perhaps needs some clarification. Imagine an experiment performed with a perfect detector baving the response function $R(x, y)=\delta(x-y)$. When accumm ulated into channels the results represent a histogram with statistical uncertainties. If we now let the channel width $\Delta$ approach an infinito esimal value dy and take the average of a large sumber of identical experiments (approaching infinity in number), then the histogram approaches a smooth function given by $T(y) d y$. In other words; $T(y)$ is proportional to the probability density function $t(y)$ of the sample emitting radiation having energy yo We can therefore write.

$$
T(y)=E N_{0} t(y) \quad, \quad 1.2
$$

where is the detector efficiency and $N_{0}$ is the average number of emissions radiated over the duration of experiment. Since $t$ (y) mast be gormalized to unity, we have

$$
\int_{0}^{\infty} T(y) d y=E N_{0}
$$

As oux wext step we can consider the effects of response $R(x, y)$ being other than a delta function. The original function $T(y)$ is modified to another function $M(x)$ according to the convolution integral

$$
M(x)=\int_{0}^{\infty} R(x, y) T(y) d t \quad I_{0} 4
$$

Note that by virtue of eq. (lo 1) the area under the function is conserved i.e.

$$
\int_{0}^{\infty} M(x) d x=\theta N_{0}
$$

If think of $x$ as the pulse beight, then the probability density function of pulse heights entering the analyzer is given by $M(x) /\left(8 N_{0}\right)$. The analyzer sorts these pulses into channels centered at positions $x_{i}$. For the moment we ignore any analyzer nonlinearity effects and assume that everywhere we have the same constant difference $x_{i+1}-x_{i}=\Delta$. If now the $i^{\text {th }}$ channel has gulse acceptance profile $P_{i}(x)$, then the contents of this channel are oa the average given by

$$
m_{i}=\int_{0}^{\infty} M(x) P_{i}(x) d x_{0} \quad 1.6
$$

As ideal channel profile would bave the shape

$$
\begin{align*}
P_{i}(x) & =\frac{1}{\Delta} \text { for }\left|x_{i}-x\right|<\frac{\Delta}{2} \\
& =0 \text { sor }\left|x_{i}-x\right|>\frac{\Delta}{2}
\end{align*}
$$

In practice the profile does not have such sharp cut-ofi edges ${ }^{\text {(1) }}$, howo ever we can use the form of (1.7) without departing too much from reality. Substituting accordingly in eq. (1.6) we obtain

$$
w_{i}=\frac{I}{\Delta} \int_{x_{1}-\Delta / 2}^{x_{i}+\Delta / 2} M(x) d x
$$

Up to now we talked about expectation values of spectra and ignored statistical effects. In a particulax experiment one observes a statistical spectrum which we denote by $m_{i}$. Throughout this thesis primen will be used exclusively fox the purpose of indicating statistical quantities As shown in Appendixs I the probability of obtaixing a particular vaiue $m_{i}$ is given by the Poisson Irequency fuxctiox

$$
p\left(m_{i}^{\prime}, m_{i}\right)=\frac{m_{i}^{\prime}}{\left(m_{i}^{\prime}\right)!} e^{-m_{i}}
$$

The same appendix also shows that for large $m_{i}$ expression (1.9) is closeIy approximated by a Gaussian frequency function, i.e. we have the asymptotic expression

$$
p\left(m_{i}^{0}, m_{i}\right) \sim \frac{1}{\sqrt{2 \pi m_{i}}} e e^{\frac{\left(m_{i}^{0}-m_{i}\right)^{2}}{2 m_{i}}}
$$

We car thus separate the statistical part of the spectrum and write
where

$$
\begin{align*}
& m_{i}^{\prime} \sim m_{i}+s_{i}^{\prime} \\
& s_{i}^{\prime}=m_{i}^{\prime}-m_{i}^{\prime}
\end{align*}
$$

and the statistical part has the erequeacy function

$$
f\left(s_{i}^{\prime}, m_{i}\right)=\frac{1}{\sqrt{2 \pi m_{i}}} e^{\frac{\left(s_{i}^{\prime}\right)^{2}}{2 m_{i}}}
$$

Collecting the results of eqs. ( 1.4 ) , ( 1.8 ) and ( 1.11 ) we have the finel mathematical model of the observed spectrum,

$$
m_{i}^{0} \approx \frac{1}{\Delta} \int_{x_{i}-\frac{\Delta}{2}}^{x_{i}+\frac{\Delta}{2}} d x \int_{0}^{\infty} R(x, y) T(y) d y+s_{i}^{i}
$$

This spectrum can be modified somewhat by addition of puises which occur within the electronic resolving time. Such events may be due to chance or they may occur on account of two nuclear transitions being in prompt cascade.

Corrections (both experimental and analytical) for these effects have been discussed by various authors $(2,3,4,5)$ and a short summary is given in Appendiy II. In the main Eext we shall assume that, if necessary, the appropriate corrections have meer made and that eq. (1.14) represents the experimental spectrumo

Our problem is to determine $T(y)$ as accurately as possible from the given set of observed values mio An exact determination is of course impossible, mainly due to the statistical uncertainty in $s_{i}$ and the integration of data into a finite number of channels.

There appear to be two main avenues of approach in determining T(y) o The first approach is based on inversion of the response matrix and will be treated in Chapter II. In this approach no assumptions about the particular form of $T(y)$ are necessary; however the condition that $T(y)$ be non-negative can be put to effective use in peducing statistical uncertaino ties. The second approach discussed in Chapter III requires that $T(y)$ be described by a model function in which the parameters are adjusted to give a best fit to the data. Usually this approach results in a weighted least squares calculatione

Chapter II includes a detailed description of matrix inversion methods applied to resolution corrections of $\beta$-ray spectra taken with organm ic scintillators. Linear Fermi-plots are obtaiaed for almost the entire sange of the energy spectruxe In addition this chapter presents a derive ation of an inverse matrix in closed form for a certain class of infinite size response matrices. The class is limited to response functions which Lave shapes sssentially independent of energy.

Chapter III includes discussion of two-dimensional time-correlation
experiments. A method of reducing the large scale problem to a simplified solution by parts is presented together with the statistical properties of the parameters thus obtained. A model calculation with $\gamma$-ray coincidence spectra is used to illustrate how one may determine $\gamma$-ray cascades, branching ratios and absolute transition rates independently of knowledge of detection efficiencies. The effects of various least squares calculation weighting functions are also studied.

The important class of least squares calculations with model functions which are non-linear in their parameters is discussed in some detail in Chapter III. Conditions are given under which the probability density function of estimates of parameters departs significantly from a Gaussian shape. Approximate probability density functions are derived for a certain class of models and compared to the results of Monte Carlo calculations.

Chapter IV summarizes the results and discusses relative merits of the two above-mentioned approaches in determining $T(y)$. It is shown how under certain conditions the two approaches merge to become identical.

## CHAPTER II

RESPONSE MATRIX APPROACH

### 2.1 Matrix Model of System Response

In this chapter we shall treat the detector response problem by methods of matrix algebra. The experimental spectrum $m_{i}^{\prime}$ of model equation (1.14) is already of a form which can be expressed by a vertically arranged vector (one column matrix)


As indicated by the subscript of the last component, we are limiting ourselves to consideration of a data field with N channels. In a similar way we can define the random vector

$$
\frac{16 n}{2^{n} \cdots \cdots w^{n} \cdot N^{n} \cdot 1^{n} \cdot}
$$

The response function $R(x, y)$ and true spectrum $T(y)$ have in principle continuous arguments and cannot be exactly represented by matrices with a finite number of elements. It thus becomes necessary to make the approximating assumption that the true spectrum $T(y)$ can be replaced by a digitized version
where

$$
\begin{aligned}
& T(y) \longrightarrow \sum_{j=1}^{N}{ }^{N}{ }_{j} \delta\left(y-y_{j}\right)_{\theta} \quad 2.1 .3 \\
& \text { B}_{\text {d }}=\frac{2}{\Delta} \int_{y_{j}-\frac{\Delta}{2}}^{y_{j}^{+\frac{\Delta}{2}}} T(y) d y . \quad 2.1 .4
\end{aligned}
$$

Note that these definitions preserve the original normalization condition:

$$
\begin{equation*}
\int_{0}^{\infty} T(y) d y=\sum_{i=1}^{N} \quad t_{j}=\varepsilon N_{0} \tag{2. 105}
\end{equation*}
$$

Substitution of definitions (2.1.3) and (2.1.4) in model equation (1. 1 4) yields the result

$$
m_{i}^{g}=\sum_{j=1}^{N} \frac{1}{\Delta} \int_{x_{i}-\frac{\Delta}{2}}^{x_{i}+\frac{\Delta}{2}} R\left(x_{, y_{j}}\right) d x t_{j}+s_{i}^{0} \quad 2.1 .6
$$

We can now define a response matrix

where the elements are given by

$$
r_{i j}=\frac{I}{\Delta} \int_{x_{i}-\frac{\Delta}{2}}^{x_{i}+\frac{\Delta}{2}} \cdot R\left(x_{o} y_{j}\right) d x_{0}
$$

Using these definitions we cass replace eq. (2.1.6) by the simplified expression

$$
a_{i}^{i}=\sum_{j=1}^{N} \quad r_{i j} t_{j}+s_{i}^{i}
$$

By the rules of matrix multiplication and addition this last statement is just the matrix equation

$$
M^{g}=R E+S^{\theta}
$$

where $T$ is a vertical vector with components $t_{j}{ }^{\circ}$
Solution for $T$ can be attempted by methods of matrix algebra. If one obtains the inverse of response matrix $R$, then multiplication of eq. (2.1.10) by $\mathrm{R}^{-1}$ from the left will produce the result

$$
\underline{R}^{-1} \underline{M}^{0}=\underline{I}+\underline{R}^{-1} \underline{S}^{0}
$$

Uafortunately the right hand side of this expression contains the undesirable statistical vector $R^{-1} \underline{S}^{\prime}$. For inverse matrices $R^{-1}$ containing elements much larger than unity this vector may represent a great magnification of the original statistical uncertainties, even to the exteat as to overshadow the structure of the true spectrum F. This is notably true for Gaussian response shapes having widths extende ing over many channels. However, there are cases where $R^{-1}$ does not contain large elements and practical application of matrix inversion
can be made, as is illustrated by the response corrections to $\beta$-ray spectra treated in a later section of this chapter. Even in cases where magnification of statistical uncertainties is large the results might be salraged by applying the condition that the true vector $T$ must be non-negative. This procedure may result in considerable reduction of statistical deviations. The statistical effects created by various inverse matrices $\mathrm{R}^{-1}$ and the application of non-negativity are discussed in the last two sections of the present chapter.

Expression (2.1.3) contains the assumption that the true spectsum consists of delta functions placed in the middle of each channel., Consequently a spectrum $T(y)$ containing entries at intermediate posita ions has no representation in the matrix model. When $\mathrm{T}(\mathrm{y})$ is a slowly varying function of $Y$, this limitation is not serious since the set of "sampling" values $t_{j}$ adequately defines $T(y)$. However, if the true $T(y)$ contains a delta function between channel centers, the unfolded spectrum $\mathbb{R}^{-1} M^{0}$ tends to split the contents between the two adjacent channels. Therefore the matrix model cannot distinguish one delta function placed between channel centers from two delta functions placed at ceaters of adjacent channels, This effect is illustrated by calculations in section 2.5.

A number of numerical methods of unfolding (correction for response matrix) have been previously reported $(6,7,8,9,10)$; These methods generally require involved calculations, the difficulty of which increases rapidly with $N$, the number of channels used. Apparently the simplest to apply are the iterative methods $(6,7)$ : However, their rate of convergence depends considerably on the type of response matrix $\mathrm{R}_{\boldsymbol{p}}$

In particular, when the response has a Gaussian shape, the convergence can be very slow, The Gaussian resolution correction has been treated analytically by Dixon et al ${ }^{(11)}$ : Their solution for $T(y)$ is given in series form containing increasing orders of derivatives of the observed spectrum $M(y)$., Since high order derivatives are difficult to evaluate the series must perforce be terminated. This has the effect of introducing oscillations in the calculated spectrum. Various methods of numerical matrix inversion are discussed in section 2.3.

Since in practice the inversion of matrices by numerical metbods is limited to rank not much larger than, say, 50 , the present work is concentrated on attempts to obtain a representation of the inverse matrix in closed form. A great advantage of this approach is the fact that data fields of almost unlimited number of channels can be considered. Computer calculations can be performed quite rapidly without the dangex of large rounding-off errors that are often encountered during numerical inversion of large size matriceso
2. 2 Correction for Analyzer Nonlinearity

Usually analyzers possess some degree of nonlinearity which results in pulse beights being distributed into bins of unequal width Do, Although as a rule slight, the effects of this phemomenon are twofold: firstly, the abscissa,or pulse-height scale is distorted and secondly, channels with greater widths receive more than their proper share of counts. For linear pulse-height scales the response matrix can often be derived from physical considerations (see section 2.4 ox $\beta$-ray spectra). It may thus be desirable to correct the observed spectrum by transferring its contents to a new set of channels which span equal pulse-beight intervals.

The required pulse-height vs. channel curve can be obtained by the use of a calibrated pulser. From the set of axalyzer addresses $\boldsymbol{x}_{9}$ obtained at various pulsemeight settings $y$, the following curve was derived for a 256 channel analyzer:

$$
y=A\left(B+x+1.09 \times 10^{-4} x^{2}\right)
$$

The parameters $A$ and B represent the gain and origin shift respective= Iy. Thus for a particular set-up of instruments only these two parameters need be determined for a full calibration.

Having obtained the explicit relationship $y=f(x)$ we can proceed with the correction as illustrated in Figure 2. Channel bins which vary in width on the pulse-height scale are shown at the top. The position of the right edge of the $i^{\text {th }}$ bin is calculated by subw stituting $x=i+\frac{l}{2}$ in formula (2.2.1). The new bins of constant width $\Delta y$ are shown at the bottom. Counts from the channel bins are transferred to the bottom bias in direct proportion to the correspo onding overlap. If counts in the original spectrum are giwen by $n_{i}$, then accordiag to Figure 2 the counts in the $j^{\text {th }}$ bin of the new spectrum become

$$
m_{j}=\frac{y_{i} \cdots(j-1) \Delta y}{J_{i}-J_{i-1}} n_{i}+\frac{j \Delta y-y_{i}}{y_{i+1}-y_{i}} n_{i+1} . \quad \sqrt{2} .2 .2
$$

In this derivation it was assumed that all pulses spanned by the width of an original channel occurred with equal probability. Whea the spectrum has a pronounced slope in the region considered this assumption is not strictly true since the rate of pulse arrival differs at the two channel sides. However, if this effect is too pronounced,

FIGURE 2
CORRECTION FOR ANALYZER NONLINEARITY

mach arailable information is lost due to channel integration and the need for a new experiment with finer chennel mesh would be indicated.

### 2.3 Numerical Methods of Matrix Inversion

This section is not essential to the main development presented in this thesis since the methods proposed are largely dependent on response matrices which can be inverted in closed form: However, for the sake of completeness ${ }_{0}$ it is deemed advisable to include a short discussion of some numerical inversion methods. A short study of inc version difficulty as a function of matrix size is also provided., This latter consideration will perhaps serve to illustrate the advantages gained by having the inverse matrix in closed forme,

In this section it is being assumed that the inverse matrix $\mathrm{R}^{-1}$ exists. A necessary and sufficient condition for this to be true is that matrix $R$ have rank $N$ (i.e. all $N$ columns of the matrix constitute a set of linearly independent $N$-dimensional vectors). An equivalent necessary and sufficient condition is that $\operatorname{det}(\underline{R}) \neq 0$. The response matrices encountered here generally satissiy these conditions so that their inverses can be calculated. Section 2.5 will cover a few special types of response matrices which under certain conditions become singuo las.

The subject of matrix inversion has received attention in numer. ous volumes (see for example references $(12,13,14,15,16)$ ). A number of papers $(27,18,19)$ mave discussed the subject from the point of view of electronic computer application.

Numerical methods of matrix inversion can be divided into two broad classes o direct methods and iterative methods. The direct methods
obtain the required solution in a finite number of operations, whereas iterative methods only approach the true solution with each iteration. After a certain number of iterations the remaining corrections may become insignificant and the process is then terminated. For large size matrices the iterative methods are often more economical in effort and may yield more accurate answers.

Perhaps the simplest of direct methods is the Gauss ${ }^{\circ}$ method of systematic elimination. The process is analogous to solution of linear equations by progressive elimination of unknown variables. One wishes to find the inverse matrix ${\underset{R}{ }}^{-1}$ such that

$$
\underline{R}^{-\mathbb{Z}} \underline{R}=I
$$

where $I \mathbb{I}$ is an $N$ by $N$ size identity matrix baving elements equal to unity along the principal diagonal and zero elsewhere. Suppose that we divide the matrices $\mathbb{R}^{-1}$ and $I$ into $N$ row vectors and denote the resultm ing vectors by $\mathrm{R}_{\mathrm{i}}^{-1}$ and $\mathrm{I}_{\mathrm{i}}$ respectively. Then for each value of index $i$ we have a set of $\mathbb{N}$ linear simultaneous equations

$$
\underline{R}_{i}^{-1} \quad R=I_{i}
$$

which cas be written out in the detailed form

$$
\begin{aligned}
& r_{i 1}^{-1} r_{11}+r_{i 2}^{-1} r_{21}+0 .+r_{i N}^{-1} r_{N 1}=\delta_{i 1} \\
& r_{i 1}^{-1} r_{12}+r_{i 2}^{-1} r_{22}+\ldots+r_{i N}^{-1} r_{N 2}=\delta_{i 2} \\
& \ldots \\
& r_{i 1}^{-1} r_{1 N}+r_{i 2}^{-1} r_{2 N}+\ldots . .0+r_{i N}^{-1} r_{N N}=\delta_{i N} .
\end{aligned}
$$

The symbol $\delta_{i j}$ is a Kroenecker delta having the property

$$
\begin{align*}
\delta_{i j} & =0 \text { for } i \neq j, \\
& =1 \text { for } i=j .
\end{align*}
$$

As an initial step the first equation of group (2.3.3) is divided by $r_{11^{\circ}}$, Next it is multiplied by $r_{12}$ and subtracted from the second equation, which gives us a new second equation with the unknown variable $r_{i l}^{-1}$ eliminated. In a similar way we can eliminate $r_{i l}^{-1}$ from all succeeding equations, which concludes the first stage., In the next stage we divide the new second equation by the coefficient of $r_{i 2}^{-1}$ and then proceed to eliminate this variable from all succeeding equations using the same process as described in stage one . After $N$ stages of such calculations we end up with a triangular set of equations.

$$
\begin{align*}
r_{i 1}^{-1}+r_{i 2}^{-1} a_{21} & +r_{i 3}^{-1} a_{31}+\ldots+r_{i N}^{-1} a_{N 1}
\end{align*}=b_{i 1}, ~ r_{i 2}^{-1}+r_{i 3}^{-1} a_{32}+\ldots+r_{i N}^{-1} a_{N 2}=b_{i 2} .
$$

$$
r_{i N}^{-1}=b_{i N}
$$

The coefficients $b_{i k}$ are zero for $k<i$ since up to the $i^{\text {th }}$ stage there is no contribution from the Kroenecker deltas.

It is interesting to note the number of multiplications and divisions performed: These operations require a greater effort than subtractions and hence provide an indication of the over-all difficulty. For the moment we shall consider only the operations performed on the left side of the equality sign. In the $k^{\text {th }}$ stage we are operating on ( $N-k+1$ ) equations. There are ( $N-k$ ) numerical divisions by the coefficient of $r_{i k}^{-1}$. The subsequent eliminations will require ( $N-k$ ) multiplications performed for each of the ( $N-k$ ) equations. Thus the total
number of operations performed in the $k^{\text {th }}$ stage is

$$
M_{k}=(N-k)+(N-k)^{2}
$$

Sumaing the operations of all stages we obtain the result

$$
\begin{aligned}
M_{T} & =\sum_{k=1}^{N} M_{k} \\
& =\frac{N}{3}\left(N^{2}-1\right)_{0}
\end{aligned}
$$

These $M_{T}$ operations define all coefficients $a_{\text {max }}$ which need not be recalculated.

Calculation of coefficients $b_{i j}$ on the right side of equality signs will require additional operations the number of which will depend on index io Consider the $k^{\text {th }}$ stage of elimination, where $k \geqslant i$. There will be one division and ( $N-k$ ) multiplications. Summing over all stages with $k \geqslant i$ we get the number of operations

$$
\begin{aligned}
N_{i} & =\sum_{k=i}^{N}(N-k+1) \\
& =\frac{i^{2}}{2}-\left(N+\frac{3}{2}\right) i+(N+1)\left(\frac{N}{2}+1\right)
\end{aligned}
$$

To obtain the full inverse of matrix $R$ we shall need to obtain a triaso gular form akin to (2.3.5) for each value of index i. Therefore we sum the $N_{i}{ }^{\text {s }} \mathrm{s}$ with the result

$$
\begin{align*}
N_{T} & =\sum_{i=1}^{N} N_{i} \\
& =\mathbb{N}\left(\frac{N^{2}}{6}+\frac{N}{2}+\frac{1}{3}\right)
\end{align*}
$$

Hence the total number of multiplications and divisions required to obtain the $N$ triangular sets (2.3.5) is given by

$$
\begin{align*}
T & =M_{T}+N_{T} \\
& =\frac{N^{2}}{2}(N+1)
\end{align*}
$$

The final calculation of inverse matrix elements is performs ed by a process of back-substitution. In equations (2.3.5) the element $x_{i k}^{-1}$ is given by

$$
r_{i k}^{-I}=b_{i k}-\sum_{\ell=k+1}^{N} r_{i \ell}^{-1} a_{\ell k} .
$$

which requires ( $N-k$ ) multiplications. For all values of $k$ the number of maltiplications is

$$
\begin{aligned}
B_{i} & =\sum_{k=1}^{N}(N-k) \\
& =\frac{N}{2}(N-1) .
\end{aligned}
$$

This number is the same for all N sets of equations (2.3.5), so that the total backasubstitution eifort can be represented by

$$
B=\frac{N^{2}}{2}(N=1)
$$

We are now in a position to write down the overall effort factor for complete matrix inversion

$$
\begin{align*}
E & =T+B \\
& =\mathbb{N}^{3}
\end{align*}
$$

This cubic relationship presents a serious practical limitation on the size of matrices which can be inverted by the direct methods.

In the foregoing derivation it was assumed that the divisor used
in the $k^{\text {th }}$ stage of elimination was non-zero. When this condition does not hold it is necessary to use a modified method. Even when the divisor is finite but small the results may be unsatisfactory due to accumulation of large rounding-off errors. A way out of these difficulties is to search each equation for the largest coefficient and then use it as the divisor" for the purpose of elimination. This method is sometimes referred to as the "Gauss' Method with Selection of the Pivotal Element"。

Another method popular in automatic computer calculations is the Gauss-Jordan method of elimination. During the $k^{\text {th }}$ stage of elimination the unknown $r_{i k}^{-1}$ is eliminated from all the preceding as well as the succeeding equations. Thus the equations (2.3.3) are reduced to diagonal form and no back-substitution is necessary. However a larger number of operations is required during the process of elimination with the result that the total number of multiplications and divisions remains $N^{3}$.

There are various other direct methods of inversion which shall not be discussed here. In particular, when the matrix is symmetrical there are special techniques available which provide a greater economy in effort. These techniques are amply discussed in the literature and the interested reader may find them in the references quoted above.

The special case of a triangular matrix shall be considered since some of the response matrices in this thesis take that form. Consider the upper-triangular matrix R with elements $\quad r_{i j}=0$ for $i>j$. In this case the matrix equation (2.3.1) can be represented by the set of algebraic equations

$$
\sum_{i=1}^{j} r_{k i}^{-1} r_{i j}=\delta_{k j}
$$

Note that in equation (2.3.11) when $k>j$ we also have $k>i$. Therefore the right hand side is zero and the equation is satisfied by setting

$$
r_{k i}^{-1}=0 \quad \text { for } k>i
$$

In other words, the inverse matrix is also upper-triangular and equation (2.3.11) can be replaced by

$$
\sum_{i=k}^{j} r_{k i}^{-1} r_{i j}=\delta_{k j}
$$

Triangular matrices are particularly easy to invert since the process can be carried out row by row independently of other rows. Consider a general $k^{\text {th }}$ row of the inverse matrix. Using expression (2.3.12) we can write a set of ( $N-k+1$ ) equations which already are in triangular form

$$
(j=N) r_{k k}^{-1} r_{k N}+r_{k k+1}^{-1} r_{k+1 N}+\ldots+r_{k N}^{-1} r_{N N}=0
$$

$(i=k+1) r_{k k}^{-1} r_{k k+1}+r_{k k+1}^{-1} r_{k+1} k+1 \quad=0$ $(j=k) r_{k k}^{-1} r_{k k} \quad=1$.

Thus the inverse elements of $k^{\text {th }}$ row can be calculated by a simple process of back-substitution. Evaluation of $r_{k k}^{-1}$ will require one division. The next term $r_{k k+1}^{-1}$ will require one multiplication and one division. It is easily seen that the number of operations increases by
one in each step. Hence the total number of operations required is given by

$$
\begin{aligned}
B_{k} & =\sum_{\ell=1}^{(N-k+1)} \ell \\
& =\frac{k^{2}}{2}-\left(N+\frac{3}{2}\right) k+\frac{N}{2}(N+3)+1
\end{aligned}
$$

The number of operations for all rows of the inverse matrix is obtained by summing over $k$, i.e.

$$
\begin{align*}
B & =\sum_{k=1}^{N} B_{k} \\
& =N\left(\frac{N^{2}}{6}+\frac{N}{2}+\frac{1}{3}\right)
\end{align*}
$$

For large values of $N$ this effort factor is about one sixth of the factor for general matrix inversion.

In many response matrices the diagonal elements are large in magnitude compared to other elements. Under these conditions the iterative methods may have good convergence properties (15) and lead to fairly rapid calculation of the inverse matrix. An example of iterative techniques is provided by the Gauss-Seidel method discussed below.

Equations (2.3.3) can be rewritten in the following form:

$$
\begin{align*}
& r_{i 1}^{-1}=\frac{1}{r_{11}}\left(\delta_{i 1}-r_{i 2}^{-1} r_{21}-\ldots-r_{i N}^{-1} r_{N 1}\right) \\
& r_{i 2}^{-1}=\frac{1}{r_{22}}\left(\delta_{i 2}-r_{i 1}^{-1} r_{12}-\ldots-r_{i N}^{-1} r_{N 2}\right)
\end{align*}
$$

$$
r_{i N}^{-1}=\frac{1}{r_{N N:}}\left(\delta_{i N}-r_{i 1}^{-1} r_{12}-\ldots-r_{i N-1}^{-1} r_{N-1 N}\right)
$$

One starts with initial guesses for the unknows $r_{i j}^{-1}$ and substitutes them at the right hand sides of equations (2.3.15). The resultant values at the left are the first approximations $\left(r_{i j}^{-I}\right) 1_{1}$ In the second step these new vaiues can be substituted at the right sides, yielding second approximations $\left(r_{i j}^{-1}\right)_{2}$ This process is repeated until the changes become insignificant. The last estimates $\left(r_{i j}^{-1}\right)_{n}$ give the $i^{\text {th }}$ row elements of the inverse matrix.

### 2.4 Application to $\beta$-Ray Spectra Obtained with Organic Scintillators

One of the first attempts to correct $\beta$ aroy spectra for resolution effects was made by Freedman et al ${ }^{(6)}$. Their studies revealed that the response to monoenergetic electrons was of the form shown in Figure 3 where the peak is associated with the totally absorbed $\beta$-rays。 In addition to the peak there is a constantwheight tail which arises from partial energy losses due to scattering from the crystal (20).

The method proposed by Freedman et al ${ }^{(6)}$, was an iterative proceduree One makes a first estimate ${ }^{[1}$ for the true spectrum and thes obtaias a second estimate $\mathbb{T}_{2}$, from the relation

$$
\underline{T}_{2}=T_{1}+M^{\prime \prime}-\mathbb{T}_{-1}
$$

where $M^{\circ}$ is the observed spectrum and $\underset{\sim}{R}$ is the response matrix. This process is contixued using the algorithm

$$
T_{m+1}=T_{m}+M^{\theta}-R T_{m}
$$

By using progressive back-substitution we can express this last equation in terms of the first estimate $T_{1}$. The result is

Here $I$ is the identity matrix and the inverse $\mathbb{R}^{-1}$ is assumed to exist.

$\operatorname{If} \lim _{m \rightarrow \infty}(\underline{I}-\underline{R})^{m}=O_{s}$
then $\lim _{n \rightarrow \infty} \quad T_{-m+1}=R^{-1} M^{\prime}$.
and successive estimates approach the unfolded spectrum. Freedman et ${ }^{\text {al }}{ }^{(6)}$ found that for a good initial estimate only two to four itera ations were required to obtain a spectrum $T_{m+1}$ such that $R T_{T+1}$ reproduced the observed spectrux $M^{8}$ within about $2 \%$

This iterative procedure has the advantage that the inverse matrix $\mathrm{R}^{-1}$ does not have to be evaluated directly. As indicated in the last section, the inversion of matrices having sizes corresponding to data fields of 200 or more channels can be quite difficult. Even if the full response matrix could be ixverted there are reasons why its use might be undesixable. The response shapes shown in Figure 3 contain the Gaussiass resolution function. As will be shown in section 2.5 , corsection for this effect involves an inverse matrix with very large positive and negative elements. Consequently the statistical uncertainties oxiginally present in $M^{8}$ can be grossly magnified.

When the iterative formula (2.4.2) is used to correct for the Gaussian response alone, the statistical deviations show a roughly linear increase with each iteration. If the results finally converge to the rem quired unfolded spectrum then this expansion of statistical deviations can be expected to continue until a levelling-off is reached near values corresponding to multiplication by the inverse matrix of Gaussias response. It is therefore important to obtain a very good first estimate which would allow the iterations to be kept dows to a minimam.

The method of response correction presented here was first desm cribed in a paper by Slavinskas, Kennett and Prestwich ${ }^{(21)}$. It is
based on writing the full response matrix as a product of an uppertriangular matrix and a Gaussian matrix. The triangular matrix represents the tail part of the response and can be inverted in closed form. After correction for this response effect is made, the Fermioplot of the $\beta_{0}$ ray spectrum is found to be very nearly linear. This allows a very good first guess $\mathbb{T}_{1}$ of the true spectrum in formula (2.4.1). Thus only one iteration is required to correct for the remaining Gaussian resolution It will be shown later in this section that multiplication by the inverse of the triangular matrix produces only a small effect on the statistical deviations。

The response functions show in Figure 3 can be represented by the convolution of a Gaussian $G\left(E_{9} E_{0}\right)$ with a second function $L\left(E_{9} E_{0}\right)_{8}$ namely

$$
R\left(E_{8} E_{0}\right)=\int L\left(E_{9} E_{1}\right) G\left(E_{1}, E_{0}\right) d E_{1}, \quad 2.4 .4
$$

where

$$
G\left(E_{,} E_{0}\right)=\frac{1}{\sqrt{2 \pi} \sigma_{0}} \exp \left[\frac{\left(E-E_{0}\right)^{2}}{2 \sigma_{0}^{2}}\right]
$$

2.4 .5 $L(\sigma, 8)$
and

$$
\begin{array}{rlr}
I\left(E_{9} E_{0}\right) & =\frac{k_{0}+\delta\left(E-E_{0}\right)}{k_{0} E_{0}+1} & \text { for } E \leqslant E_{0} \\
& =0 & \text { for } E>E_{0}
\end{array}
$$

Parameter $k_{0}$ is related to the tail height of response to $\beta$ rays with energy $\mathrm{E}_{0}$ and $\sigma_{0}$ is the standard deviation of the Gaussian resolution at the same energy (full width at half-maximum $\approx 2.35 \sigma_{0}$ ), All three functions are gormalized to have unit area.

Using the digitizing approximations discussed in section 2.1 we can replace the response functions by matrices of rank $N_{0} \quad L\left(E_{9} E_{0}\right)$ is replaced by an upper-triangular matrix $I$ in which the rows are indexed by values of $E$ and the columas indexed by $E_{0^{\circ}}$ Similarly we can replace $G\left(E, E_{0}\right)$ by a matrix $G_{\text {. }}$ The overall response matrix $\underline{R}$ can then be writo tes in the form

$$
\underline{R}=\underline{I} G \cdot \quad 2.4 .7
$$

In accordance with equation (2.1.10) we can express the obo
served spectrum as

$$
M^{0}=\underset{\sim}{G} \underset{\underline{G}}{ }+S^{0} 9 \quad 2.4 .8
$$

where $\mathbb{T}$ is the true spectrum without response effects and $\underline{S}$ is a random vector of statistical deviations. Operation on $M^{8}$ by the inverse matrix $\underline{I}^{-1}$ will result io

$$
G T+L^{-1} S^{8}
$$

Apart from statistical deviations this is just the true spectrum $T$ as "seen" through a Gaussian resolution. Since $\beta-r a y$ spectra are continuous and usually have widths large compared to that of Gaussian resolution, the spectrura $\mathcal{E} T$ is very similas to the true spectrum $T_{0}$

Evaluation of inverse matrix $\mathrm{L}^{-1}$ is straight-forward and can be written in closed form. This is a great advantage since we are no longe er limited to small arrays and one can obtain the inverse for almost any size data field. The delta function in equation (2.4.6) represents $\beta$-rays which deposit their full energy within the crystal. If we define the area under this full-energy peak by $\alpha_{j}$ we can write

$$
\alpha_{j}=\frac{I}{k_{j}{ }_{j}+I}
$$

The area under the tail is given by the complement

$$
\beta_{j}=1-\alpha_{j} .
$$

With these definitions matrix I has the form


Derivation of the inverse matrix is given in Appendix III. If we define the quantity

$$
\gamma_{j}=I-\frac{\beta_{j}}{(j-I) \alpha_{j}}
$$

where $j>1$, then the inverse matrix elements are given by

$$
\begin{aligned}
& l_{i i}^{-1}=\frac{1}{\alpha_{i}}, \\
& \ell_{i j}^{-1}=0 \quad \text { for } \quad i \quad j \quad j \\
& d_{i j}^{-1}=\frac{1}{\alpha_{i}}\left(\gamma_{j}-1\right) \prod_{k=i+1}^{j-1} \gamma_{k} \text { for } \quad i<j .13
\end{aligned}
$$

Trial and error calculations with experimental spectra having end-point energies up to 2.27 MeV indicated that linear Fermi-plots can be obtained by using constant values for parameters $\alpha_{j}$ and $\beta_{j}$. This fact implies that the fraction of electrons scattered outside
the crystal is independent of incident energy. Experimental support for this conclusion is provided by the work of Bothe ${ }^{(22)}$ who found that for a given scattering material the back-scattering coefficient of low-energy electrons is constant. The values used in the succeeding calculations are $\alpha=0.72$ and $\beta=0.28$.

From equation (2.4.12) it is apparent that all values of $\gamma_{j}$ are less than unity. This means that in any given row of $\underline{L}^{-1}$ the diagonal element $I / \alpha=1.39$ has the largest numerical value and all elements to the right progressively become smaller. Therefore the magnification of statistical uncertainties produced by the operator $I^{-1}$ is small enough as to be of no practical significance.

Sample calculations of resolution correction were performed on three essentially single group spectra of ${ }^{90},{ }^{91},{ }^{143} \operatorname{Pr}$ and a mixture of ${ }^{90} \mathrm{Sr}$ with 90 Y in secular equilibrium. All experimental. spectra were obtained under identical conditions with the sources mounted 1 cm from a $5 \times 5 \mathrm{~cm}$ type $N \mathrm{D}-102$ organic scintillator coupled to an EMI 9536 photo-multiplier tube. The, pulses were fed to a DD2 amplifier and recorded in a multi-channel analyzer. Sources were impregnated into filter paper and had a thickness corresponding to approximately $3 \mathrm{mg} / \mathrm{cm}^{2}$.

Before applying resolution corrections the spectra were correcta ed for analyzer nonlinearity by the use of the technique described in section 2.2. The parameters $A$ and $B$ in formula (2.2.1) were determined from internal conversion lines of ${ }^{137} \mathrm{Cs},{ }^{207} \mathrm{Bi}$ and initial estimates of the $\beta$-ray spectra end-points. All $\beta$-ray spectra were then converted from the original channel base to a linear energy scale with bins
$\Delta E=10 \mathrm{keV}$ wide. Values used for the two parameters were

$$
\begin{aligned}
\mathrm{A} & =9.444 \mathrm{keV} / \text { channel } \\
\text { and } \quad B & =12.0 \text { channels }
\end{aligned}
$$

${ }^{143}$ gr has a $\frac{5}{2}+$ ground state which decays by $\beta^{m}$ to the $\frac{7}{2}$ - ground state in ${ }^{143} \mathrm{Nd}(23)$. Thus we have a change in spin $\Delta J=1$ and a parity change, the combination of which corresponds to a first forbidden transition (24). Many first forbidden transitions have spectra correso ponding to those of allowed transitions ${ }^{(25)}$ and the case of ${ }^{143}$ Pr appears to fall in this class. The energy distribution of allowed transition $\beta$-rays is given by

$$
T(E)=D P W F\left(Z_{9} E\right)\left(E_{m}-E\right)^{2}, \quad 2.4 .14
$$

where $D$ is a constant and $F\left(Z_{9} E\right)$ is a Coulomb effect correction factor dependent on the atomic number $Z$ and electron kinetic energy $E$ o Numerical values of $F(Z, E)$ which included electron screening effects were obtained from reference ${ }^{(26)}$. The factor $F(Z, E)$ is essentially constant at high $\beta$-ray energies and has little effect on the shape of the spectrum Other factors in formula (2.4.14) are the electron momentusa

$$
p=\frac{1}{c} \sqrt{E^{2}+2 E m_{0} c^{2}}
$$

and the total electron energy

$$
W=E+m_{0} c^{2}
$$

Parameter $E_{m}$ is the total transition energy and thus represents the upper energy limit of the $\beta$-ray. When $E<E_{m}$ the remaining energy is carried away by the neutrino (nuclear recoil energy can usually be neglected).

It is customary to represent the data in the form of so-called Fermi (or Fermi-Kurie) plots. These can be obtained by plotting $\sqrt{T /(P W F)}$ as a function of $E_{\text {. The result should be a straight line }}$ with an intercegt $E=E_{n}$ on the abscissa. Thus one obtains a conven ient indication of transition energy $E_{m}$.

The $\beta^{-}$transitions of ${ }^{90}$ Sr ${ }^{90} Y$ and ${ }^{91} Y$ are characterized by $\Delta J=2$ and a change in parity. Thus they fall in the class of first forbidden transitions. On account of the spin change $\Delta J=2$ the spectra are expected to have unique shapes ${ }^{(25)}$ which are distinguished from the allowed shapes (2.4.14) by having an extra factor $S$. The shape factor was calculated from the relation

$$
S=Q^{2}+\lambda_{I} P^{2}
$$

where $Q$ is the neutrino momentum and $\lambda_{1}$ is an energy-dependent parameter (close to unity) tabulated by Kotani and Ross (27).

Langer et al ${ }^{(28)}$ have reported the necessity of an additional shape correction for 90 and ${ }^{91} Y_{Y}$. Based on their precise magnetic spectrometer measurements, this additional factor takes the form

$$
c=1+\frac{b}{w} \cdot
$$

where $w$ is the total electron energy in $m_{0} c^{2}$ units and $b$ is an empirical parameter having a numerical value between 0.2 and 0.40 Ia the calculations presented here the intermediate value $b=0.3$ was used. Due to these additional factors the expression $\sqrt{T /(\text { PWFSC ) }}$ was used to calculate Fermi plots for ${ }^{90} S_{S r}, 90_{Y}$ and ${ }^{91} Y_{Y}$.

Fermi plots of ${ }^{143} \mathrm{pr}_{8} 91_{Y}$ and 90 are shown in Figures 4,5
and 6 respectively. Each Figure contains three plots which illustrate the various stages of resolution correctiono glots labelled (a)



were calculated before application of any resolution corrections. They tend to show an upward curvature for decreasing energies. This effect is caused mainly by the low-energy tail in the scintillator response and makes the determination of end-point energy uncertain. Fermi plots (b) were obtained after multiplying the spectra by inverse matrix $\underline{L}^{-1}$. A considerable improvement in linearity is evident, so that accurate end-point determinations can now be made. The remaining slight curvature near maximum energies is presumably caused by Gaussian resolution effects. This residual has the effect of slightly modifying the apparent end-point energy. Calculations with theoretical model spectra gave an estimate of the error magnitude. A number of theoretical spectra with different maximum energies were multiplied by the Gaussian resolution matrix. Fermi plots were then calculated and "best" straight lines were drawn through the resulting points. It was found that these lines gave correct intercepts with the ordinate axis but underestimated the end-point energy by about 8 keV .

Although the two corrected intercepts thus obtained are. sufficient to determine the required experimental spectrum parameters, further calculations were applied in an attempt to correct for the Gaussian resolution. A quadratic variation of $\sigma_{0}^{2}$ with energy was assumed. After a few trial calculations the particular form adopted was

$$
\sigma_{0}^{2}=0.0027 E^{2}+E+3000
$$

where both $\sigma_{0}$ and $E$ are expressed in $k e V$. One iteration
as defined by equation (2.4.1) was applied to each of the three spectra. The first estimates $T$ were calculated by the use of corrected intercepts of Fermi plots (b). Results are shown in part (c) of Figures 4, 5 and 6. Iinearity near the end-points is seen to be considerably improved. Some points lie below the E-axis because imaginary values (obtained where the unfolded spectrum was negative) are plotted as negative numbers.

A summary of end-point energy results is given in Table I. The largest deviation from Nuclear Data Sheet values ${ }^{(29)}$ (column 2) occurs for the end-point of ${ }^{90}$ Sr, the difference being 20 keV . Errors in other isotopes are considerably smaller. This is partly caused by the fact that the experimental end-points were included in the determination of the energy scale. Since all spectra were obtained under identical experimental conditions, the energies should be internally consistent. This requirement is satisfied within the estimated uncertainty of $\pm 20 \mathrm{keV}$ 。

The two-component spectrum of ${ }^{90} \mathrm{Sr}$ and ${ }^{90} \mathrm{Y}$ was analyzed for relative intensities as well as for end-point energies. Intensities are expected to be equal since the system was in secular equilibrium with the 28 year ${ }^{90} \mathrm{Sr}$ decaying to the 64 hour ${ }^{90}$ Y. Fermi plots of the results are given in Figure 7. Part (a) shows the total spectrum after multiplication by $\underline{L}^{-1}$. The long linear part of the ${ }^{90} Y$ component provided good estimates of the intercepts. These were used to calculate a theoretical 90 response matrix and then subtracted from the total. A Fermi plot of the remaining ${ }^{90}$ Sr spectrum is shown in part (b). Finally, part (c)

## TABLE I

End-Point Energies

| Isotope | $\mathrm{E}_{\mathrm{m}}(\mathrm{MeV})$ |  |
| :---: | :---: | :---: |
|  | Nuclear Data Sheets (29) | Present Work |
| ${ }_{90}$ Sr | 0.54 | $0.56 \pm 0.02$ |
| $143$ | 0.93 | $0.93 \pm 0.02$ |
| $91_{Y}$ | 1.54 | 1.54 $\ddagger 0.02$ |
| $90 \%$ | 2.27 | $2.27 \pm 0.02$ |


shows the Fermi plot of ${ }^{90} \mathrm{Sr}$ after correction for Gaussian resolution. Experimental intercepts obtained for ${ }^{90} \mathrm{Sr}$ and 90 Y were used to calculate the corresponding theoretical spectra. These were then summed over all channels to obtain a measure of intensities. The results gave a total of ( $1.02 \pm .10$ ) $\times 10^{6}$ counts for ${ }^{90} \mathrm{Sr}$ and (1.04 $\pm 0.05$ ) $\times 10^{6}$ counts for 90 , which indicates good agreement with the expected equal intensity condition. In principle, this procedure of "stripping" could be applied to a spectrum with any number of components. The limitation, of course, is that the end-points have separations sufficiently large to leave recognizable straight Iine segments in the Fermi plot after the operation $\underline{I}^{-1}$.

It is shown in Appendix III that the number of multiplicative and divisional operations required to invert $I^{-1}$ has a quadratic dependence on matrix size $N$. This is a considerable improvement over the cubic dependence (2.3.14) which obtains for general triangular matrices. Thus one can perform rapid analysis even when the data fields are quite large. For example, using the IBM 7040 electronic computer, corrected Fermi plots were calculated in 2 min. and 16 sec. for a 64 by $64 \beta-\gamma$ coincidence spectrum.

### 2.5 Application to Energy-Independent Response Shapes

In this section we shall consider response functions of the form $R(x-y)$, which implies a fixed response shape for all incident energies $\mathbb{y}$. Nuclear radiation detectors usually do not satisfy this requirement; for example, the Gaussian resolution width in spectra obtained with $N a I(T l)$ detectors varies roughly as the square root of incident $\gamma$-ray energy. However, it is possible that future develop-
ments of analyzers ${ }^{(30)}$ will provide the option of storing the spectra on a channel base proportional to $\sqrt{\text { E. Obtaining thus an approximate- }}$ Iy constant resolution width would result in most efficient use of the available memory storage capacity (or number of channels). In some instances (e.g. lithium-drifted Ge detectors) the variation in resolution width is very small over the energy range covering one peak. With some modifications the method of this section might still be used to apply resolution corrections for such spectra.

The technique described below need not be limited to nuclear spectra alone. There is a wide class of other applications where some signal may have suffered distortion due to undesirable frequency comcomponent filtering - or equivalently convolution - effects. Such may be the case when dataare transraitted, for example, over long telephone lines. In some instances a short pulse may be purposefully distorted into a long wave train, which technique permits transmission of increased energy when the peak power of a transmitiing device is limited. Examples of this technique are found in some radar systems ${ }^{(31)}$ and in seismic exploration methods using surface sources (32). When this extended signal is received, some method of de-convolution is usually required in order to obtain the original short pulse. The present section investigates a method of demconvolution which is applicable to data having digitized form.

In the present section we shall derive a general expression for the inverse of an infinite rank response matrix $R$ having elements $r_{i j}$ which are dependent on (i-j) only. This matrix consists of identfeal columns which are shifted by one row with respect to their neigh-
bours. The derivation of inverse matrix elements will be accomplished by the use of Fourier transforms and the convolution theorem ${ }^{(33)}$.

We start by considering the convolution integral

$$
M(x)=\quad \int_{-\infty}^{\infty} R(x-y) T(y) d y, \quad 2.5 .1
$$

and then try to find an inverse function $R_{\text {in }}(x)$ such that

$$
T(x)=\int_{-\infty}^{\infty} R_{i n}(x-y) M(y) d y
$$

For this purpose we shall need the Fourier transform pair

$$
\begin{array}{ll}
M(x)=\frac{1}{2 \pi} \int_{-\infty}^{\infty} m(\omega) e^{i \omega x} d \omega \\
m(\omega)=\int_{-\infty}^{\infty} M(x) e^{-i \omega x} d x
\end{array}
$$

Capital and lower case letters shall be employed to represent the functions in the $x$-domain and the $\omega$-domain respectively. A short derivation of the convolution theorem is given in Appendix IV. Given the convolution relationship in the $x$-domain

$$
F_{3}(x)=\int_{-\infty}^{\infty} F_{1}(x-y) F_{2}(y) d y
$$

the theorem states that in the $\omega$-domain we have the product

$$
f_{3}(\omega)=\hat{f}_{1}(\omega) \hat{i}_{2}(\omega) . \quad 2.5 .4 \mathrm{~b}
$$

Similarly, for convolution in the w-domain

$$
g_{3}(\omega)=\int_{-\infty}^{\infty} g_{1}(\omega-\lambda) g_{2}(\lambda) d \lambda \quad 2.5 .5 \mathrm{a}
$$

we have

$$
G_{3}(x)=2 \pi G_{1}(x) G_{2}(x)
$$

The method of finding $R_{\text {in }}(x)$ can now be outlined as follows. Applying the convolution theorem to equations (2.5.1) and (2.5.2) we can write the products

$$
r(\omega)=r(\omega) t(\omega)
$$

and

$$
t(\omega)=r_{i n}(\omega) m(\omega)
$$

To satisfy these last two relations we mast have

$$
r_{i n}(\omega)=\frac{1}{r(\omega)}
$$

Thus the relation ( 2.5 .8 a) gives us the Fourier transform of the required inverse function $R_{i n}(x)$. Accordingly we can transform back to the $x$-domain which yields the result

$$
R_{i n}(x)=\frac{1}{2 \pi} \quad \int_{-\infty}^{\infty} \frac{e^{i \omega x}}{r(\omega)} d \omega . \quad 2.5 .8 b
$$

Often $r_{i n}(\omega)$ is unbounded and the inverse function $R_{\text {in }}(x)$ does not exist. This is usually the case with continuous functions $R(x)$. As an example we can consider the Gaussian response function

$$
R(x)=\frac{1}{\sqrt{2 \pi} \sigma} \exp \left[-\frac{x^{2}}{2 \sigma^{2}}\right]
$$

which has the Fourier transform (see Appendix IV)

$$
r(\omega)=\exp \left[-\frac{\sigma^{2} \omega^{2}}{2}\right]
$$

It is clear that the reciprocal $I / r(\omega)$ approaches infinity together with $\omega_{\text {; }}$ hence the value $R_{\text {in }}(0)$ as defined by ( 2.5 .8 b ) is infinite
and the desired inverse function does not exist. However, the inverse can frequently be found in matrix form when the response functions are digitized into discrete channel bins. The Gaussian response falls into this class of invertable digitized functions.

A form of digitization can be accomplished by making use of the following set of equally spaced delta functions

$$
D(x)=\sum_{n=-\infty}^{\infty} \delta(x-n)
$$

which, according to Appendix IV, has the Fourier transform

$$
d(\omega)=2 \pi \sum_{n=-\infty}^{\infty} \delta(\omega-2 \pi n)
$$

Thus the digitized response function can be written as the product

$$
R^{d}(x)=R(x) D(x)
$$

According to the scheme outlined above we will try to find the inverse function $R_{i n}^{d}(x)$ by calculating the Fourier transform of $R^{d}(x)$ and then transforming the reciprocal $1 / r^{d}(\omega)$ back to the $x$-domain. Proceeding with the execution of the first step we write the Fourier transiorm

$$
\begin{align*}
r^{d}(\omega) & =\int_{-\infty}^{\infty} R(x) e^{-i \omega x} \sum_{n=-\infty}^{\infty} \delta(x-n) d x \\
& =\cdots \sum_{n=-\infty}^{\infty} R(n) e^{-i \omega n} .
\end{align*}
$$

This expression contains a summation which may consist of many terms when $R(x)$ covers a wide range of integral values of $x$. For such cases it wight be more convenient to express $r^{d}(\omega)$ in another form which can be particularly useful when the function $r(\omega)$ is easily obtained.

Applying the convolution theorem to expression (2.5.11) we can write

$$
\begin{align*}
& r^{d}(\omega)=\frac{1}{2 \pi} \\
& \quad \int_{-\infty}^{\infty} r(\omega-\lambda) 2 \pi \sum_{n=-\infty}^{\infty} \delta(\lambda-2 \pi n) d \lambda \\
&=\sum_{n=-\infty}^{\infty} r(\omega-2 \pi n)
\end{align*}
$$

From expression (2.5.13) it is clear that, although $r(\omega)$ may approach zero for large values of $\omega, r^{d}(\omega)$ will not do so on account of the periodicity inherent in the sumnation. Except for special cases in which $r(2 \pi n)=0$ for all values of $n_{0}$ the reciprocal $1 / r^{d}(\omega)$ will remain finite everywhere. Figure 8 illustrates the effect of digitization by showing the reciprocal functions $I / r(\omega)$ and $1 / r^{d}(\omega)$ obtained for a Gaussian response with $\sigma=1$.

Transformation to the $x$-domain is accomplished by noting that $1 / r^{d}(\omega)$ is periodic in $\omega$ with period $2 \pi$. Thus after defining the truncated function

$$
\begin{align*}
f(\omega) & =\frac{1}{r^{d}(\omega)} & \text { for } & & |\omega|<\pi \\
& =\frac{1}{2} \frac{1}{r^{d}(\omega)} & \text { for } & & |\omega|=\pi^{\prime} \\
& =0 & \text { for } & & |\omega|>\pi
\end{align*}
$$

we can express

$$
\begin{align*}
\frac{I}{r^{\alpha}(\omega)} & =\sum_{n=-\infty}^{\infty} f(\omega-2 \pi n) \\
& =\int_{-\infty}^{\infty} f(\omega-\lambda) \sum_{n=-\infty}^{\infty} \delta(\lambda-2 \pi n) d \lambda \\
& =\frac{1}{2 \pi} \int_{-\infty}^{\infty} f(\omega-\lambda) d(\lambda) d \lambda .
\end{align*}
$$

The required inverse function $R_{\text {in }}(x)$ ts given by the Fourier transform $0 i^{1} 1 / r^{d}(\omega)$. Using (2.5.15) and the convolution theorem we can

express the inverse function in the form of the product

$$
R_{\text {in }}(x)=F(x) D(x),
$$

where

$$
F(x)=\frac{1}{2 \pi} \int_{-\pi}^{\pi} \frac{e^{i \omega x}}{r^{d}(\omega)} d \omega
$$

The inverse function (2.5.16) is seen to contain the original digitizing function $D(x)$. This factor is instrumental in replacing the de-convolution integral (2.5.2) by a summation so that the process becomes analogous to multiplication by an inverse matrix of infinite dirension. Carrying out the substitution in (2.5.2) we obtain

$$
\begin{aligned}
T(x) & =\int_{-\infty}^{\infty} F(x-y) \sum_{n=-\infty}^{\infty} \delta(x-y-n) M(y) d y \\
& ={ }_{n=-\infty}^{\infty} F(n) M(x-n)
\end{aligned}
$$

Since the response function was digitized, the spectrum $M(x-n)$ will have non-zero values only for integral values of the argument ( $x-n$ ). Thus we need consider only integral values of $x$, and upon making the transformation $m=x-n$ in equation (2.5.18) we get the result .

$$
T(x)=\sum_{m=-\infty}^{\infty} F(x-m) M(m)
$$

If the digitized response function (2.5.11) is substituted in the original convolution integral (2.5.1), we obtain

$$
M(x)=\sum_{n=-\infty}^{\infty} R(x-n) T(n) .
$$

Expressions (2.5.19) and (2.5.20) can be replaced by the matrix eqations

$$
\underline{T}=\mathrm{F} \underline{M}
$$

$$
2.5 .21
$$

and

$$
\underline{M}=\underline{R} \underline{\underline{I}} \quad \quad 2.5 .22
$$

where $F$ is an infinite square matrix having elements

$$
f_{i j}=F(i-j)
$$

and $R$ is likewise an infinite square matrix with elements

$$
r_{k \ell}=R(k-\ell) .
$$

$\underline{I}$ and $\underset{M}{M}$ are infinite one-column vectors. It is clear that

$$
E=\underline{R}^{-1} .
$$

and after collecting the results of equations (2.5.23), (2.5.17) and (2.5.13) we can write down the inverse matrix elements

$$
r_{k \ell}^{-1}=\frac{1}{2 \pi} \int_{-\pi}^{\pi} \frac{e^{i \omega(k-\ell)}}{\sum_{n=-\infty}^{\infty} r(\omega-2 \pi n)} d \omega
$$

where $r(\omega)$ is the Fourier transform of the response function $R(x)$. The denominator of the integrand can be replaced by the alternative expression (2.5.12) if the latter proves to be more convenient. The foregoing derivation used response functions which were digitized by sampling $R(x)$ at integral values of $x$ corresponding to channel positions. However, if the data are of the histogram type, it is necessary to include the effects of integration over channel width, as is done in the model (2.1.8). Using unit channel width we can write the new digitized response function

$$
R^{d}(x)=D(x) \int_{x-\frac{1}{2}}^{x+\frac{1}{2}} R(y) d y
$$

The channel profile is assumed to be flat-topped and is given by the
function

$$
\begin{array}{rlrl}
P(x) & =1 \text { for }|x|<\frac{1}{2}, \\
& =\frac{1}{2} \text { for }|x|=\frac{1}{2}, & \\
& =0 \text { for }|x|>\frac{1}{2} .
\end{array}
$$

As shown in Appendix $I V_{s}$ the Fourier transform of $P(x)$ is

$$
p(\omega)=\frac{2}{\omega} \sin \left(\frac{\omega}{2}\right)
$$

Expression (2.5.27) can be put in the form of a convolution integral

$$
R^{\alpha}(x)=D(x) \quad \int_{-\infty}^{\infty} P(x-y) R(y) d y_{2}
$$

which enables us to use the convolution theorem and obtain the Fourier transform

$$
r^{d}(\omega)=\sum_{n=-\infty}^{\infty} x(\omega-2 \pi n) \frac{\sin \frac{1}{2}(\omega-2 \pi n)}{\frac{1}{2}(\omega-2 \pi n)}
$$

Therefore the inverse matrix elements for channel-integrated response are given by

$$
r_{k l}^{-1}=\frac{1}{2 \pi} \int_{-\pi}^{\pi} \frac{e^{i \omega(k-\ell)}}{\sum_{n=-\infty}^{\infty}\left[r(\omega-2 \pi n) \frac{\sin \frac{1}{2}(\omega-2 \pi n)}{\frac{1}{2}(\omega-2 \pi n)}\right]} d \omega
$$

Some examples of matrices and their inverses are provided in Table II. An integrated expression could be obtained only for the first entry (damped sine-wave response). The integrals in other entries must be evaluated by numerical methods. In some cases the inverse matrix does not exist for certain values of the response parameters. For example, in the matrix for damped sinusoid response we must satisfy $a \geqslant \pi$ and in the matrix corresponding to a triangular response funct-


TABLE II(cont'd)
Examples of Inverse Matrices
$r_{i j} \mid$

Gaussian

$$
I_{i j}=\frac{A}{\sigma \sqrt{2 \pi}} \exp \left[\infty(i-j)^{2} /\left(2 \sigma^{2}\right)\right]
$$

$$
r_{i j}^{-1}=\frac{1}{A \pi} \int_{0}^{\pi} \frac{\exp \left(\sigma^{2} \omega^{2} / 2\right) \cos [\omega(i-j)]}{\sum_{l=\infty}^{\infty} \exp \left[2 k \pi \sigma^{2}(\omega-b \pi)\right]} d \omega
$$

ChanneloIntegrated Gaussian

$$
r_{i j}=\frac{1}{\sigma^{-} \sqrt{2 \pi}} \int_{i-j-\frac{1}{2}}^{i-j+\frac{1}{2}} \exp \left[-x^{2} /\left(2 \sigma^{2}\right)\right] d x
$$

$$
r_{i j}^{-1}=\frac{2}{\pi} \int_{0}^{\pi} \frac{\operatorname{exo}\left(\sigma^{2} \omega^{2} / 2\right) \cos [\omega(i-j)]}{\sum_{l=\infty}^{\infty}} \exp \left[2 k \pi \sigma^{2}(\omega-k \pi)\right] \frac{\sin \left[\frac{1}{i}(\omega-2 \ell)\right]}{\frac{1}{2}(\omega-2 k)} d \omega
$$

tion the parameter $q$ cannot have integral values other than unity. The last two entries contain the factor $\exp \left(\sigma^{2} \omega^{2} / 2\right)$ in the inverse matrix integrands. This may lead to very large inverse matrix elelements for high values of $\sigma$ and thus render application of the inverse matrix unpractical when the spectrum contains statistical deviations.

The matrices considered in this section have the restriction that their elements depend only on their distance ( $k-l$ ) from the principal diagonal. This restriction reflects the assumption that the shape of response functions is independent of the incident energy (e.g. constant value of $\sigma$ for Gaussian response functions). In general a response matrix has elements

$$
(j)_{r_{i j}}=(j)_{R(i-j)}
$$

Where the superscript $j$ signifies the dependence of response function on column number. Under certain conditions it is possible to obtain a good approximation to the inverse of this generalized matrix. Suppose we fix the superscript $j$ at a certain value $d$ and then obtain the inverse matrix ${ }^{(d)}{\underset{\sim}{R}}^{-1}$ by the method outlined above. This inversion can be carried out in turn for various other fixed values of $\ell$ Then we construct a matrix $\mathbb{R}_{0}^{-1}$ in which the $2^{\text {th }}$ column is given by the corresponding column of $(l)_{R^{-1}}$. If the elements of $(l) \underline{R}^{-1}$ are given by

$$
(l)_{r_{i j}}^{-1}=(l)_{F(i-j)}
$$

then the condition that

$$
R \underline{R}_{0}^{-1} \longrightarrow I
$$

can be stated in the detailed form

$$
\sum_{j}^{(j)} R(i-j) \stackrel{(k)}{F(j-k)} \rightarrow \delta_{i k}
$$

Equality would obtain in this last expression if we could replace $(j)_{R(i-j)}$ by $(k)_{R(i-j)}$. If the response shape varies siowly with energy, then close to the value of $k$ there is a range of $j$-values for which $(j)_{R(i-j)} \approx{ }^{\left.(k)_{R(i-j)}\right)}$. Suppose that outside this range of j-values the contribution to the sum in (2.5.34) is negligible; then $\underline{R}_{0}^{-1}$ can be a good approximation to the true inverse matrix $\underline{R}^{-1}$. In the case of Gaussian response with the standard deviation $\sigma(\mathbb{E})$ these necessary conditions are satisfied when $d \sigma / \partial E \ll 1$.

### 2.6 Reduction of Statistical Deviations

As was remarked in Chapter $I$, the spectra of nuclear counting experiments contain statistical deviations which have very nearly Gaussian frequency functions. Each channel of the observed spectrum therefore has an associated variance $v_{j}^{2}=m_{j}$, where $m_{j}$ is the expectation of the contents of the $j^{\text {th }}$ channel. Since the contents of the channels are mutually statistically independent, all covariances $v_{i j}^{2}$ are zero. When the observed spectrum is multiplied by an inverse matrix $R^{-1}$, the variances in the unfolded spectrum will naturally have magnitudes different from the original variances. According to the error propagation formula given in section 3.2, the new variance in channel i will be

$$
u_{i}^{2}=\sum_{j}\left[r_{i j}^{-1}\right]^{2} v_{j}^{2} .
$$

Poun this expression it is clear that when the inverse matrix $\underline{R}^{-1}$ contains elements of large magnitude, the unfolded spectrum will have greatly magnified statistical uncertainties. A lower limit on the
ratio of standard deviations in channel $i$ is given by

$$
\left(\frac{u_{i}}{v_{i}}\right)_{\min }=\left|r_{i i}^{-I}\right|
$$

Expressions (2.5.26) and (2.5.31) defining the inverse matrix elements, show that the diagonal elements always have the largest numiurical value. It is therefore of some interest to consider their magnitudes for various types of response functions and various values of response parameters. Figure 9a shows the diagonal inverse matrix element as a function of parameter $q$ for a triangular response function. The response matrix and inverse matrix are defined in the third entry of Table II, where an anplitude factor $A$ is included so as to conserve the normalization condition $\underset{j}{\sum} r_{i j}=1$. It is seen that, as parameter q takes on integral values greater than unity, the element $r_{i j}^{-1}$ becomes infinite. This has the significance that for these values of $q$ there is no unique unfolded solution. Figure $9 b$ exemplifies this fact by showing two initially different spectra which become identical after a triangular response with $q=2$ is folded in.

The inverse matrix element $r_{i i}^{-l}$ for a channel-integrated Gaussian response (see entry 5 of Table II) is shown as a function of $\sigma$ in Figure 10. The extremely rapid rise with increasing $\sigma$ is dictated by the factor $\exp \left(\sigma^{2} \omega^{2} / 2\right)$ in the integrand of the expression for inverse matrix element. Since, according to equation (2.6.2), the magnification in statistical uncertainties becomes rapidly untenable, the practical application of this inverse matrix is linited to values of $\sigma$ not much greater than unity. However, in practice it is cosirable wave a high channel density for the purpose of


(b)

$$
\text { FIGURE } 9
$$



FIOURE 10
GHANNEL-INTEGRATED
gavsslan response
preserving as much of the fine structure as possible. This leads to high values of $\sigma$ (which is measured in channel units) and we are thus faced with conflicting requirements.

Since the merits of fine channel mesh cannot be denied, the difficulty might be resolved by modifying the inverse matrix. Instead of attempting to "squeeze" the Gaussian into one channel, one may have to settle for partial unfolding which would leave some residual but narrower response function $R_{n}(x)$. The elements of a "partial inversion" matrix (see Appendix IV) are given by

$$
r_{k \ell}^{-1}=\frac{1}{2 \pi} \int_{-\pi}^{\pi} \frac{e^{i \omega(k-l)}}{r^{d}(\omega)} \sum_{n=-\infty}^{\infty} I_{n}(\omega-2 \pi n) d \omega .
$$

where $r^{\alpha}(\omega)$ is defined by $(2.5 .12)$ and (2.5.13). For a Gaussian response the value of $1 / r^{d}(\omega)$ is especially large near the integration limits $\omega= \pm \pi$. Therefore, if we are to keep the magnitude of $r_{k l}^{-1}$ down to relatively low values, we must choose the residual $R_{n}(x)$ in such a way that its Fourier transform $r_{n}(\omega)$ is very small. or zero near $\omega= \pm \pi$. It may also be desirable that $R_{n}(x)$ have no negative values. This can be ensured by choosing a $r_{n}(\omega)$ such that it is the convolution of some other function with itself. According to the convolution theorem $R_{n}(x)$ is then some function squared and hence cannot have negative values. Within these requirements there is considerable room for variation in the particular form of $R_{n}(x)$ chosen. Some future effort may be warranted in search of an optirrum shape such that for a given reduction in resolution width the magnification of statistical uncertainties is minimized.

A powerful method of reducing statistical deviations can be
derived by making use of the a priori knowledge that intensity spectra cannot have negative values. If the observed spectrum vector is $M^{\prime \prime}$ (as defined in (2.1.10)), then, due to the random nature of $\underline{S}^{\prime}$, the unfolded spectrum $\underline{R}^{-1} \underline{M}^{\prime}$ will generally contain some negative as well as positive components. The condition of non-negativity can be imposed by adding a vector of corrections $\underline{C}$ to $\mathbb{M}^{8}$ such that the result gives

$$
\underline{R}^{-1}\left(\underline{M}^{\prime}+\underline{C}\right) \geqslant 0
$$

Equation (2.6.4) defines an infinite set of allowed vectors © - Out of this set we select a particular vector so as to maximize the likelihood function

$$
I=\left[\begin{array}{ll}
\prod_{i} & \frac{1}{\sqrt{2 \pi} v_{i}}
\end{array}\right] \exp \left[\begin{array}{ccc}
-\frac{1}{2} & \frac{c_{i}^{2}}{i} & \frac{v_{i}^{2}}{2}
\end{array}\right]
$$

where $c_{i}$ are the components of vector $\underset{C}{ }$ and $v_{i}^{2}$ are the variances in spectrum $M^{\prime \prime}$. Maximum $L$ is obtained by minimizing the exponent, i.e. we have the condition

$$
\sum_{i} \frac{c_{i}^{2}}{v_{i}^{2}} \longrightarrow \min
$$

It will be shown in section 3.4 that good estimates of the weights $I / v_{i}^{2}$ are provided by $I /\left(m_{i}^{\prime}+1\right)$, where $m_{i}^{\prime}$ are the components of the observed spectrum M".

The problem as stated above can be given a geometrical interpretation. If we consider spectra limited to N channels, then expression (2.6.4) represents $N$ Iinear inequalities with $N$ independent variables $c_{i}$. Each inequality defines a region to one side of a hyper-plane in N-dimensional space. All inequalities combined define a feasible region for the vector $\underline{C}$ in this space. The solution
is given by the particular point in this feasible region which will minimize the objective function $\sum_{i}\left(c_{i}{ }^{2} / v_{i}{ }^{2}\right)$. This is a problem in mathematical programming for which there are several methods of so(34)
lution available. However, the difficulty of solution is comparable to numerical matrix inversion and in the following sample calculation we shall use another approach which is simpleminded but perhaps, for this application, more practical.

An example of unfolding by matrix inversion is given in
Figure 11. Part (a) shows the assumed true spectrum $T(x)$ in the form of a number of delta functions. Note that $T(x)$ includes two delta functions at the non-integral positions 71.5 and 92.5 for which there is no representation in the matrix model, as explained in section 2.1. Part (b) shows the spectrum $M^{\prime \prime}(x)$ obtained by folding in a channelintegrated Gaussian response with $\sigma=1$ and adding appropriate statistical deviations. The deviations were obtained by sampling from a set of normally distributed random numbers. After multiplication by the inverse matrix $\underline{R}^{-1}$ we obtain the spectrum shown in part (c). Large positive and negative oscillations caused by magnification of statistical uncertainties are seen to obscure the true spectrum almost completely. Part (d) shows the result after application of non-negativity calculations, which were based on a very rapid iterative procedure. In each iteration a search was made for the largest negative value in the spectrum, which was found, say, at channel $y$. Then the following combination of inverse functions $F(x)$ was added to the spectrum:

$$
A F(x-y)-\frac{A}{2}[F(x-y-1)+F(x-y+1)]
$$



The amplitude $A$ was chosen so as to make the spectrum component at channel $y$ equal to zero. Inclusion of the last two terms avoids the introduction of a positive bias. The iterations were continued by searching again for the largest negative component and repeating the same process as above. If after $n$ iterations the spectrum is $T_{n}(x)$, then the addition of expression $(2.6 .7)$ to $T_{n}(x)$ affects only three channels in the spectrum $R \mathbb{T}_{n}$. The value $A$ is added to channel $y$ and $\frac{A}{2}$ is subtracted from the channels $(y-1)$ and $(y+1)$.

The spectrum in Figure $11(d)$ was obtained after 1200 such iterations performed on a base of 200 channels. The result is a considerable improvenent over the spectrum in part (c). All doublets, except the one near channel 70 are now clearly resolved. It is interesting to note that the delta function originally placed at position 92.5 ended up distributed between channels 92 and 93 in about equal proportions. In principle, the resuits should be better if the constraint of non-negativity were applied properly by minimizing the objective function (2.6.6).

### 3.1 Model Equation and Solution for Parameters

In Chapter II we dealt with the matrix model in which it was assumed that the true spectrum can be represented by the components of a vector T. Various methods of response correction were suggested for the purpose of obtaining a spectrum which would approximate $T$ as closely as possible. An advantage associated with these wthods is that no previous knowledge of the mathematical form of I is required. If the form of response matrix $\underline{R}$ is known, then the same set of inversion calculations should be applicable to all spectra, independently of the number or intensity of components present. However, as was shown in section $2.1_{2}$ the matrix model is limited in that some types of spectra cannot be exactiy represented by the vector ${ }^{T}$. For instance, ${ }^{\text {P }}$ cannot inciude transitions having energies which do not correspond to channel mid-points. This limitation is particularly inportant when it is desired to make accurate transition energy determinations.

The least squares approach does not have this limitation, since it permits consideration of any energy positions. In fact. the energies can be used as continuously variable least squares solution paraueters. However, the least squares approach requires the ise of a model function wich involves the response functions and a parametrized form of an assumed true spectrum. Construction
of this model function usually requires a detailed inspection of the observed spectrum followed by an educated guess as to the number and positions of the energy transitions present.

Application of least squares methods to the analysis of onedimensional muclear spectra has received considerable attention in previous literatuce ( $36,37, \ldots, 48$ ) In the case of NaI(TL) $\gamma$-ray spectra one great practical difficulty is encountered in the generation of a library of accurate response functions. The usual procedure $(47,48)$ is to observe the response shapes of a number of standard isotopes having the simplest possible spectra, preferably consisting of a single energy transition. These response shapes are divided into segments which can be fitted by some convenient parametrized functions. The parameters thus obtained are in turn fitted to some appropriate functions of incident $\gamma$ rray energy. By methods such as these one obtains a set of parameters which permit calculation of response functions $R\left(E_{0} I_{0}\right)$ for a continuous range of incident energies $\mathrm{E}_{0^{\circ}}$ Since pariticular details of these methods are given elsewhere ${ }^{(47,48)}$, they will not be discussed in this thesis and the subsequent development will proceed with the assumption that the response surface $R\left(E_{9} E_{0}\right)$ is knowno

The surface $R\left(x_{0} y\right)$ is assumed here to be a continuous function of both its arguments. In a nuclear $\gamma$-ray spectrum we need consider only a set of $n$ discrete $\gamma$-rays having energies $y_{i}$, where $i=\ldots, 28 \ldots$, n. Each $\gamma$-ray $\mathfrak{i l l}$ have an associated response Iunction which is digitized by integration into channel bins. If asch channel is assigned unit widtho then these response functions
have the form

$$
X_{i}(x)=\int_{x-\frac{1}{2}}^{x+\frac{1}{2}} R\left(x_{1}, y_{i}\right) d x_{1}
$$

where x is allowed to have only integral values. Summing over the contributions of all $\gamma$-rays we can write down the observed spectrum model Aunction

$$
m_{x}=\sum_{i=1}^{n} a_{i} x_{i}(x)
$$

The parameters $a_{i}$ are indicative of the $\gamma$-ray intensities and can be expressed by

$$
a_{i}=\varepsilon_{i} \mathscr{H}_{i}
$$

where $\mathscr{N}_{i}$ is the number of transitions $\gamma_{i}$ occurring in observation time $T$, and $\varepsilon_{i}$ is the detection efficiency of transition $\gamma_{i}$ e

According to the statistical model derived in Chapter $I_{0}$ the observed spectrum $m_{x}^{\prime}$ contains statistical deviations $s_{x}^{\prime \prime}$ and can be written as the sum

$$
m_{x}^{0}=m_{x}+s_{x}^{0}
$$

For large numerical values of $m_{x}$ the frequency function of $s_{x}^{\circ}$ can be assumed to be normal with zero mean and variance $m_{x}$ This frequency function is given by $f\left(s_{x^{\prime}}^{n} m_{x}\right)$ in the expression (1.1.3)。 Since the $N$ channels under consideration are statistically independent, one can obtain an N-dimensional frequency function for all the deviations $s^{9}$ by simply writing the product

$$
L=\prod_{x=1}^{N} f\left(s_{x^{9} m_{x}}^{\ell}\right)
$$

$$
=\text { (const). } \exp \left\{-\frac{1}{2} \sum_{x=1}^{N} \frac{1}{n_{x}}\left[m_{i x}^{0}-\sum_{i=1}^{n} a_{i} x_{i}(x)\right]^{2}\right\} \cdot \quad 3.105
$$

I is known as the likelihood function of parameters $a_{i}$ and finds extensive application to statistical estimation problems.

We can estimate the coefficients $a_{i}$ by the method of "maximum likelihood". It will be shown in the following section that this method, based on maximizing $L_{9}$ is particularly advantageous in this application since it provides unbiased estimates with minimum possidle variances. Obviously, $L$ is maximized by minimizing the exponent in expression (3.1.5). This requirement leads directly to the well known condition of weighted least squares, namely

$$
\sum_{x=1}^{N} \frac{1}{m_{x}}\left[m_{x}^{0}-\sum_{i=1}^{n} a_{i} x_{i}(x)\right]^{2} \longrightarrow \min
$$

Differentiating (3.1.6) with respect to $a_{k}$ and equating the result to zero yields a normal equations

$$
\sum_{x=1}^{N} \frac{m_{x}^{q} x_{k}(x)}{m_{x}}=\sum_{i=1}^{n} a_{i}^{q} \sum_{x=1}^{N} \frac{x_{i}(x) x_{k}(x)}{m_{x}}
$$

where $k=I_{2} 2, \ldots, n_{0}$ If the weights $1 / m_{x}$ were known, then we would be in position to use equations (3.1.7) for solution of the n unknown quantities $a_{i}^{\prime}$. In section 304 it will be shown that the expression $I /\left(n_{x}^{8}+1\right)$ provides good estimates of the weighting factors; thus for all practical purposes we may assume that the weights are known and proceed with the solution for parameters $a_{i}$ 。 Since equations (3.1.7) are lInear, they can be conveniently represented in matrix notation. We start by defining the N-row and n-column matrix
3.1 .8
3.1 .9
3.1 .10
and

$$
A=\left\|\begin{array}{l}
a_{1} \\
a_{2} \\
0 \\
0 \\
0 \\
a_{n}
\end{array}\right\|
$$



$$
\underline{M}=\underline{X} A
$$

The statistical observed vector $\underline{M}^{\circ}$ was already defined in (2.1.1). Since the least squares solution for A will have statistical properties (indicated by the use of a prime), we define the solution vector

$$
A^{0}=\left\|\begin{array}{c}
a_{1}^{0} \\
0 \\
a_{2} \\
0 \\
0 \\
0 \\
a_{12}^{0}
\end{array}\right\|
$$

We now can write the vector of residuals
where

$$
\begin{align*}
& D^{8}=M^{0}-M^{8 \%} 0 \\
& M^{88}=X A^{0}
\end{align*}
$$

$$
3.1 .13
$$

is the least squares estimate of vector M Opon introduction of the diagonal matrix of weights
the condition of least squares $(3,1,6)$ can be rewritten

$$
\underline{D}^{9^{T}} \underline{Q} \underline{D}^{0} \rightarrow \min _{0}
$$

$$
3.1 .16
$$

where the superscript $T$ indicates the transposition of vector $D^{\circ}$ 。 Finally, we are in position to rewrite the normal equations (3.1.7) in the new form

$$
E^{T} W D^{0}=0
$$

If we define the design matrix

$$
B=X^{T} W X .
$$

then equation (3.1.17) becomes

$$
\underline{B} \underline{A}^{0}=\underline{X}^{T} \underline{W} \underline{M}^{\circ} .
$$

To obtain a solution for $A^{\prime}$ we can maltiply this last equation by $\underline{B}^{-1}$, provided the inverse matrix exists. The result is

$$
A^{8}=\underline{B}^{-1} \underline{X}^{T} W \underline{M}^{8}
$$

Matrix $B$ has some important properties which we shall now investigate. First of all, it is easily seen that $B$ is symetric about the principal diagonal. In fact, from the definition (3.1.18) we have

$$
\underline{B}^{T}=\left(\underline{X}^{T} W \underline{X}\right)^{T}=\underline{X}^{T} \underline{W} \underline{X}=\underline{B}
$$

It is shown in textbooks on statistics (e.g. Cramér (49) that. provided the columns of matrix $X$ form a set of $n$ linearly independent vectors (i.e. $\underline{X}$ has rank $n$ ), then $B$ is a positive definite matrix. This means that $\operatorname{det}(B)>0$ and that the inverse matrix $\underline{B}^{-1}$ exists. The condition that response functions $X_{i}(x)$ be linearly independent should always be satisfied in a properly set up model. Iinear dependence indicates that one response function can be written as a linear combination of others and is therefore redundant. We can therefore assume that the inverse $\underline{B}^{-1}$ exists and the solution (3.1.20) is obtainable. Since the inverse of a symmetric matrix is also symmetric, we have

$$
\left(\underline{B}^{-1}\right)^{T}=\underline{B}^{-1}
$$

## Qstistical Properties of the Least Squares Solution

The last section dealt with a method of obtaining estimates for the coefficients $a_{i}$ and showed that the required solution exists.

In matrix notation this solution was expressed by equation（3．1．20）． Now we turn to the investigation of statistical properties of the solution vector $A$ 。

First we shall show that the estimate $A^{8}$ is unbiased，i．e． the expectation in $A^{\circ}$ or，in other words，the average value that would be obtained after a large number（approaching infinity）of ideritical experiments and solutions，is given by Ao Using the symbol E for expectation we can write

$$
\begin{align*}
\mathbb{E}\left(\underline{A}^{8}\right) & =\mathbb{E}\left(\underline{B}^{-1} \underline{X}^{T} \underline{\underline{M}}\right) \\
& =\underline{B}^{-1} \underline{X}^{T} \underline{E}\left(\underline{M}^{0}\right) .
\end{align*}
$$

Matrices $\underline{B}^{-1}$ ．$\underline{X}^{T}$ and $[$ were moved to the left of the expectation sign。 since they are definite quantities．Since $M^{0}$ is assumed to be normal－ Iy distributed with mean value $M_{0}$ we have

$$
E\left(\underline{M}^{q}\right)=\underline{M}=\underline{X} .
$$

Thus equation（3．2．1）can be rewritten to yield the result

$$
E\left(\underline{A}^{\theta}\right)=\underline{B}^{-1} \underline{B} \underline{A}=A .
$$

which proves the assertion that estimates $A^{\circ}$ are unbiasedo
The statistical properties of a normally distributed vector， such as $M^{\prime}$ can be conveaiently summarized by the use of a covariance matrix ${\underset{M}{M r}}$ ．Individual elements of this matrix are defined by the expression

$$
\left.\left\{c_{M 0}\right\}_{i j}=\mathbb{E} \Gamma\left(m_{i}^{0}-m_{i}\right)\left(m_{j}^{0}-m_{j}\right)\right] \quad 0.2 .4
$$

According to this definition the diagonal elements represent var－ iances of individual components $m_{i}^{0}$ 。 whereas the off－diagonal elements are covariances between different components $m_{i}^{0}$ and $n_{j}^{0}$ 。 In the
particular case ofobserved vector $M^{9}$. the components are statistically independent and hence all covariances are zero. Since the weight matrix $\underset{\sim}{f}$ is diagonal and has diagonal elements equal to reciprocals of variances $m_{i}$, we immediately have the result

$$
C_{W i}=\underline{U}^{-1}
$$

The covariances in vector $A$ are generally not zero and we shall now derive the full covariance matrix ${\underset{A}{A}}^{\circ}$. If we make the substitution

$$
\underline{P}=\underline{B}^{-1} \underline{X}^{T} \underline{W}
$$

then we can rewrite equation (3.1.20) in the simple form

$$
A^{\circ}=\underline{P} M^{0}
$$

There is a theorem (50) which states that, for a normally distributed vector $\underline{M}^{\prime}$, the vector $A^{\prime \prime}$ is also normally distributed and has the covariance matrix

$$
\underline{C}_{A}=\underline{P} \underline{C}_{M} \cdot \underline{P}^{\mathrm{T}}
$$

Accordingly we can write

$$
\begin{align*}
\underline{C}_{A_{i}} & =\underline{B}^{-1} \underline{x}^{T} \underline{W}^{-1} \underline{W}^{T} X \underline{B}^{-1} \\
& =\underline{B}^{-1} .
\end{align*}
$$

Equation (3.2.8) and the expectation statement (3.2.3) completely define the statistical properties of solution vector $A^{\circ}$ 。

It is possible to sev that the least squares solution A"provides a best unbiased estimate of the true vector $A$ in the sense that components $a_{s}^{\circ}$ have miniman possible variances. For this purpose we construct Fisher's information insx 프 which is defined ${ }^{(51)}$ to have
elements

$$
h_{i j}=-\mathbb{E}\left(\frac{\partial^{2} \log I}{\partial a_{i} \partial a_{j}}\right)
$$

Carrying out this double differentiation on the exponent of likelihood function $I$ we obtain the elemenus

$$
h_{i j}=\sum_{x=1}^{N} \frac{X_{j}(x) X_{j}(x)}{M_{x}}
$$

which are identical to the elements of design matrix Bo We now let A have some other unbiased estimator $A^{\prime \prime}$ with a corresponding covariance matrix $\mathcal{G}_{\text {An }}$. Then according to an irportant theorem ${ }^{(52)}$ we have

$$
V^{T} \underline{C}_{A V} \underline{V} \geqslant \underline{V}^{T} \underline{H}^{-I} V
$$

where $V$ is any a-dimensional vertical vector. Since $H=\underline{B}$ and. consequently, $H^{-1}=\underline{C}_{A}$, we can write the inequality

$$
\underline{V}^{T} \underline{C}_{A} \underline{V} \geqslant \underline{V}^{T \mathrm{I}} \underline{C}_{A} \quad \underline{V}
$$

If we choose a particular vector $V$ in which all but the $j^{\text {th }}$ component are zero, the above inequality recluces to

$$
\left\{c_{A B}\right\}_{j j} \geqslant\left\{c_{A 0}\right\}_{j j}
$$

This last expression proves that the estimate $A^{\circ}$ has minimum possible variances.

In some applications it is necessary to deal with a nonlinear function of the components $a_{i}^{\prime}$. When the variances of $a_{i}^{\prime}$ can be assumed to be small, it is possible to dexive approximate expressions for the statistical properties of the function $\approx^{\circ}\left(a_{1}^{0}, a_{2}^{0}, \ldots, a_{n}^{n}\right)$ o Using a Taylor series expansion, in which quadratic and higher order
terms are neglected, we can wite

$$
f^{\gamma}-f=\frac{2}{2} \frac{\partial x}{\partial a_{i}}\left(a_{i}^{0}-a_{i}\right)
$$

Squaring both sides and taking theis expectation values we obtain

$$
\begin{aligned}
E\left(\hat{i}^{9}-\hat{L}\right)^{2} & =\sigma_{i}^{2} \\
& =\sum_{i_{2}}^{n}=1 \frac{\partial \hat{I}}{\partial a_{i}} \frac{\partial \dot{L}}{\partial a_{j}}\left[E\left(a_{i}^{0}-a_{i}\right)\left(a_{j}^{0}-a_{j}\right)\right] \cdot \quad 3.2 .14
\end{aligned}
$$

The quantity within square brackets is, by definition, the covariance between components $a_{i}^{0}$ and $a_{j}^{0}$ (or variance if $i=j$ ) and is equal to an element of the covariance matrix $C_{A}$. Hence we can write down the variance of function $\stackrel{Y}{ }^{9}$ 。

$$
\sigma_{i}^{0}=\sum_{i_{0} j=1}^{n} \frac{\partial I}{\partial a_{i}} \frac{\partial i}{\partial a_{j}}\left\{C_{A}\right\}_{i j}
$$

The expectation in $f^{0}$ is given by $I=f\left(a_{1}, a_{2} \ldots a_{n}\right)$ and within the range of values where the linear assumption (3.2.13) is good, the probability density function of $f^{\prime}$ is

$$
P\left(\dot{r}^{8}\right)=\frac{1}{\sqrt{2 \pi} \sigma_{\hat{y}}^{0}} \exp \left[-\frac{\left(\dot{S}^{8}-\hat{y}\right)^{2}}{2 \sigma_{\hat{y}}^{2}}\right]
$$

A commonly used measure of "goodness of fiti is the function

$$
\begin{aligned}
X^{2} & =\underline{D}^{N} \underline{W}^{N} \\
& =\sum_{x=1}^{N} \frac{I}{m_{x}}\left[m_{x}^{0}-\sum_{i=1}^{n} a_{i}^{0} x_{i}(x)\right]^{2} \cdot 3.2 .17
\end{aligned}
$$

The excectation in $X^{2}$ is equal to the number of degrees of freeComs ( $N-\mathrm{N}$ ), and the degree of departure from this value provides a measure of considence in the correctness of the least squares model
employed. Excessively low values would indicate that the estinates of variances in the original data were too high, whereas too large values of $\mathcal{K}^{2}$ might be due to the model (3.1.2) being incorrect. The tolerable limits for values of $\chi^{0^{2}}$ can be found in tables of most textbooks on statistics.

### 3.3 Application to Two-Dimensional Coincidence Spectra

In this section we shall consider the least-squares analysis of two-dimensional time correlation experiments performed with detectors $x$ and $y$. Two methods of solution for the coincidence com efficients will be discussed. The first method is based on a full least squares solution which in some instances may require the inversion of an impractically large design matrix. The second method avoids this difficulty by breaking up the model equation into parts and then performing a least squares calculation for each part in turn. In general, the two solutions lead to estimates having different statistical properties. Full covariance matrices will be derived for both methods of solution and comparisons of relative variances will be made.

Some practical uses of the coincidence coefficients $a_{i j}$ will be also discussed together with some sample calculations. It will be shown how they can be used to determine $\gamma$-ray cascades, branching ratios and even absolute decay rates without requiring knowledge of detector efficiencies.

Suppose that detector $x$ produces response functions $X_{i}(x)$ 。 where $i=\eta_{2}, \ldots, n_{x}$. The subscript $i$ indexes the $n_{x}$ types of nuclear transitions to which detector x is sensitive. Similarly we
let detector $y$ have $n_{y}$ response functions $Y_{j}(y)$. In general the numbers $n_{x}$ and $n_{y}$ may be different. We could have, say, detector $x$ sensitive to $\beta$-rays and detector $y$ to $\gamma$-rays. The function $X_{i}(x)$ would in this case represent the experimental $\beta$-ray spectrum with end-point $E_{i}$, whereas $Z_{j}(y)$ would be the response to a $\gamma$-ray of energy $\mathbb{E}_{j}$. In the particular case of a $\gamma-\gamma$ coincidence spectrum we expect to have $n_{x}=n_{y}$. We assume that the normalization condition $\sum_{X} X_{i}(x)=\sum_{y} Y_{j}(y)=1$ applies, With this stipulation the response functions can be thought of as being frequency functions over their respective channel bases.

If counters $x$ and $y$ operate freely without coincidence gating we have the one-dimensional model spectra

$$
M_{x}(x)=\sum_{i=1}^{n_{x}} a_{x i} x_{i}(x)
$$

and

$$
M_{y}(y)=\sum_{j=I}^{n_{y}} a_{y j} Y_{j}(y)
$$

where the coefficients can be expressed. in the form

$$
\begin{array}{ll}
a_{x i}=\varepsilon_{x i} \mathscr{N}_{i}, & 3.3 .3 \\
a_{y j}=\varepsilon_{y j} \mathscr{N}_{j} . & 3.3 .4
\end{array}
$$

The number of nuclear transitions $\gamma_{i}$ in time $T$ is given by $\mathscr{N}_{i}$ and the total detection efficiencies are represented by $\mathcal{E}_{\mathrm{xi}}$ and $\mathcal{E}_{\mathrm{yj}}{ }^{\circ}$

When coincidences between detectors $x$ and $y$ are demanded the resultant spectrum fills a two -dimensional array of size $N_{x} N_{y}$. Since detector $x$ is independent of detector $y_{0}$ the two-dimensional frequentcy function for a coincidence $(i, j)$ is simply the product $X_{i}(x) X_{j}(y)$ 。

Upon summing the coincidences between all possible pairs ( $i, j$ ) we obtain the spectrum model

$$
M_{2}\left(x_{\imath} y\right)=\sum_{i=1}^{\sum_{j=1}^{2}} \sum_{i j}^{Z_{y}} a_{i}(x) Y_{j}(y)
$$

The coincidence coefficients are given by

$$
a_{i j}=\varepsilon_{x i} \varepsilon_{y j}\left[\mathbb{W}_{i j}(\theta) \mathscr{N}_{i j}+k \mathscr{N}_{i} \mathscr{N}_{j}\right], \quad 3.3 .6
$$

where $W_{i j}(\theta)$ is the angular correlation function averaged over all relative angles subtended by the two detectors, $\mathcal{N}_{i j}$ is the number of coincident transitions ( $i, j$ ) and $k$ is related to the coincidence resolving time $\tau$ through the expression

$$
2 \tau=k T
$$

The second term in equation (3.3.6) represents the contribution due to chance coincidences.

The function $\vec{W}_{i j}(\theta)$ is usually unknown during the initial stages of a decay-scheme study. However, its variations are relativeIy small and for most practical purposes we can assume (53)

$$
\vec{W}_{i j j}(\theta) \approx 1 . \quad 3.3 .8
$$

Solution for coefficients $a_{i j}$ can be effected by the application of least squares techniques to model equation (3.3.5). There are $n_{x} n_{y}$ coefficients; hence the full least squares solution requires the inversion of an $n_{x} n_{y}$ by $n_{x} n_{y}$ size design matrix Clearly, this matrix can become prohibitively large when $n_{x}$ and $n_{y}$ reach the neighbourhood of 10 or more.

For large scale poblems we can use a solution by parts in which the dimensions of x and y are treated separately. Model equat-
ion (3.3.5) can be rewritten in the form of two coupled equations (1)

$$
\begin{align*}
& M_{2}\left(x_{y} y\right)=\sum_{i=i}^{n_{x}} q_{i}(y) X_{i}(x) \\
& q_{i}(y)=\sum_{j=i}^{n_{y}} a_{i j} Y_{j}(y) .
\end{align*}
$$

If the value of $y$ is fixed, equation $(3.3 .9)$ can be used as a model for a least squares fit in the x-dimension. This calculation, yielding $n_{x}$ coefficients $q_{i}(y)$, can be carried out for all the $N_{y}$ values of $y$ in the data field. As a result, least squares estimates are obtained for all $n_{x} N y$ intermediate coefficients $q_{i}(y)$. Next we can fix the value of $i$ and, using model equation (3.3.10), perform a least squares fit in the $y$-dimension. After $n_{x}$ such calculations one obtains estimates of all $n_{x} n_{y}$ values of $a_{i j}$ 。

The solution by parts requires a total of $N_{y}+n_{x}$ least squares calculations, each involving the inversion of a size $n_{x}$ by $n_{x}$ or $n_{y}$ by $n_{y}$ design matrix which is considerably smaller than the $n_{x} n_{y}$ by $n_{x} n_{y}$ design matrix in the full solution. This reduction in matrix size is a decided advantage in that a larger number of nuclear transitions can be considered in the analysiso

There is another practical advantage to the solution by parts. The simultaneous solution does not provide estimates of intermediate coefficients $q_{i}(y)$ which represent a one-dimensional spectrum in coincidence with the $i^{\text {th }}$ muclear transition. These spectra are important because they provide additional insight into the model. For instance, they may provide a basis for improving the accuracy of empioyed response functions or they may even reveal the existence of
new transitions which were not included in the original model. When we write down definite (non-statistical) quantities, the coupled equations (3.3.9) and (3.3.10) are mathematically equivalent to the model equation (3.3.5). In the least squares calculations we are dealing with statistical intermediate coefficients $q_{i}^{\prime}(y)$ which in general do not satisfy an equation like (3.3.10) but are distributed around its values. Therefore the solution by parts can be expected to yield estimates of $a_{i j}$ somewhat different from the true least squares estimates of the full solution. The character of this difference shall be investigated below. It shall be shown that both types of estinates are unbiased and that the solution by parts gives estimate variances greater than or equal to the fiull solution variances. This reduction in accuracy is the price exacted for the practical advantages offered in the solution by parts. Fortunately it turns out that the increase in variances is usually small - only a Lew percent in the examples considered below. Full covariance matrices of the coefficients $a_{i j}$ will be derived for both methods of solution.

As was done in the foregoing sections, we assume that the observed spectrum $M_{2}^{8}(x, y)$ is normally distributed with the variance in each channel being $M_{2}\left(X_{0} y\right)$. The contents of different channels are assumed to be statistically independent. We could again write down a likelihood function and maximize it with respect to the parameters. However we go directly to the ensuing least squares condition

$$
\underset{x=1}{\sum_{x=1}^{N_{x}}}{\underset{y=1}{N}}_{N_{y}}^{y}\left[M_{2}^{\prime}(x, y)-\sum_{i=1}^{n_{x}} \sum_{j=1}^{n_{y}} a_{i j} x_{i}(x) y_{j}(y)\right]^{2} / M_{2}(x, y) \rightarrow \min
$$

Differentiating (3.3.11) with respect to $a_{k \ell}$ and equating the result to zero we obtain $n_{x} n_{y}$ normal equations

$$
\sum_{x=1}^{N_{x}} \sum_{\sum=1}^{N_{y}} \quad x_{k}(x) Y_{2}(y) M_{2}^{0}(x, y) / M_{2}(x, y)
$$

$=\sum_{i=1}^{n_{x}} \sum_{j=1}^{n_{y}} a_{i j} \sum_{x=1}^{N_{z}} \sum_{y=1}^{N_{y}} x_{k}(x) Y_{L}(y) x_{i}(x) Y_{j}(y) / M_{2}\left(x_{8} y\right) \quad$ : 3.3 .12 where $k=1,2, \ldots, n_{x}$ and $\ell=1,2, \ldots, n_{y}$.

Expression (3.3.12) represents the normal equations of the full least squares solution. In order to discuss the solution and its properties in a systematic manner, it is convenient to use methods of matrix algebra. Some matrices will consist of smaller submatrices and their detailed expression would normally require cumbersome spatial arrangements of elements. For this reason a special type of notation is adopted. Matrices and vectors will be represented by enclosing a general element between double vertical bars. Two superscripts outside the bars will indicate the number of rows and columns. Curly braces will be used to denote a general matrix element, wheter it be just a number or a submatrix. For example, we define the matrix of line shapes

$$
\begin{align*}
& \underline{Y}=\left\|\begin{array}{llll}
y_{1}(1) & y_{2}(1) & \ldots & y_{n_{y}}(1) \\
y_{1}(2) & y_{2}(2) & \ldots & r_{n_{y}}(2) \\
\ldots & \ldots & \ldots & \ldots
\end{array}\right\| \\
& y_{1}\left(N_{y}\right) \\
& y_{2}\left(N_{y}\right) \ldots \\
& y_{n_{y}}\left(N_{y}\right)
\end{align*} \|
$$

The subscripts $y$ and $j$ indicate the row and column of the general. element. Carrying this notation one step further we define a matrix of submatrices

$$
\underline{z}=\left\|\left\{\left\|\left\{x_{i}(x) y_{j}(y)\right\}_{y, j}\right\|^{y^{N}, n_{y}}\right\}_{x_{0} i}\right\| \begin{array}{ll}
N_{x^{\prime}, n_{x}} & \\
& 3.3 .14
\end{array}
$$

The outer superscripts $N_{x}$ and $n_{X}$ refer to the number of rows and columns of subratrices. For elements the number of rows and columns is $N_{x} N_{y}$ and $n_{X} n_{y}$ respectively.

The spectrum $M_{2}\left(x_{0} y\right)$ can be represented by the vertically arranged vector (one column matrix)

$$
\underline{M}_{2}=\left\|\left\{\left\|\left\{M_{2}(x, y)\right\}_{y, I}\right\|^{N_{y^{\prime}}^{\prime}}\right\}_{x, I}\right\|^{N_{x^{9}} I}
$$

Likewise, for the coefficients $a_{i j}$ we use the vector

$$
A_{2}=\left\|\left\{\left\|\left\{a_{i j}\right\}_{j, I}\right\|^{n_{y}, I}\right\}_{i, I}\right\|^{I_{x}, I} . \quad 3.3 .16
$$

With these definitions we can write model equation (3.3.5) tin the simplified matrix notation

$$
\underline{M}_{2}=\underline{Z} A_{2} .
$$

To rewrite the normal equations (3.3.12) we make use of two further definitions. Let $\underline{D}_{2}^{8}$ be the vector of residuals

$$
\underline{D}_{2}^{8}=\underline{M}_{2}^{0}-Z \mathbb{A}_{2}^{9}
$$

where the primes on $\mathrm{M}_{2}$ and $\mathrm{A}_{2}$ indicate the coresponding statistical vectors. We also introduce the diagonal matrix of weights

$$
\underline{W}_{2}=\left\|\left\{\left\|\left\{\frac{\delta_{y j}}{M_{2}(x, y)}\right\}_{y, j}\right\|^{N_{y} N_{y}} \quad \delta_{x i}\right\}\right\|_{x, i} \cdot 3.3 .19
$$

Now the normal equations (3.3.12) can be replaced by the equivalent expression

$$
\underline{B}_{2} \underline{A}_{2}^{0}=\underline{Z}^{T} \underline{W}_{2} \underline{M}_{2}^{0}
$$

where

$$
\underline{B}_{2}=\underline{2}^{T} \underline{W}_{2} \underline{2}
$$

is the symmetric design matrix. If we obtain the inverse of this matrix, then we can write down the full least squares solution

$$
\underline{A}_{2}^{8}=\underline{B}_{2}^{-I} \quad \underline{Z}^{T} \underline{G}_{2} \underline{M}_{2}^{8} .
$$

A sufficient condition for the existence of inverse matrix $\mathrm{B}_{2}^{-1}$ is that the response functions $X_{i}(x)$ and $Y_{j}(y)$ form two sets of linearly independent vectors (see Appendix V).

The same arguments as were used in section 3.2 can be applied here to derive the statistical properties of solution vector $A_{2}{ }^{\circ}$. We have the obvious result that the normally distributed estimate $A_{2}^{\prime}$ is unbiased and that its covariance matrix is given by

$$
\underline{C}_{A_{2}^{0}}=B_{2}^{-1}
$$

Since $A_{2}^{8}$ is the result of a full least squares solution, its individual components have minimura possible variances.

We now turn to the problem of the solution by parts and consideration of its statistical properties. As stated before, this solution is expected to yield estimates $A_{2}^{\ddot{m}}$ different from the full solution estimates $A_{2}^{\prime}$. It will be necessary to introduce the following quantities: the matrix of response functions in the x-dimension

$$
\underline{x}=\left\|\left\{x_{i}(x)\right\}_{x_{g} i}\right\|^{N_{x^{n}} n_{x}}
$$

the data vectors

$$
M_{y}^{\prime}=\left\|\left\{M_{2}^{\prime}\left(x_{0} y\right)\right\}_{x_{2} I}\right\|^{N_{x^{\prime}} I}
$$

vectors of intermediate coefficients

$$
Q_{y}^{0}=\left\|\left\{q_{i}^{0}(y)\right\}_{i, I}\right\|^{e^{n}, I} .
$$

the diagonal matrix of weights

$$
\begin{align*}
& \underline{v}_{y}=\left\|\left\{\frac{\delta_{x j}}{M_{2}\left(x_{s} y\right)}\right\}_{x_{0} j}\right\|^{N_{x}, N_{x}},
\end{align*}
$$

and the positive definite design matrix

$$
\underline{F}_{y}=\underline{x}^{T} \underline{H}_{y} \underline{x} .
$$

Using model equation (3.3.9) we can obtain least squares estimates of the intermediate coefficients

$$
Q_{y}^{0}=E_{y}^{-1} \underline{X}^{T} \underline{W}_{y} M_{y}^{0} .
$$

It can be easily shown that the estimates $Q^{Q}$ are unbiased and that they have the covariance matrices

$$
\mathrm{C}_{Q_{y}^{s}}=F_{F_{y}}^{-1}
$$

Next we turn to model equation (3.3.10) which we use for fitting the data in the yodimension In the set of $n_{x} N_{y}$ coefficients $q_{2}^{0}(y)$ we will now have to fix the value of $i$ and let $y$ vary. Therefore the coefficients must be rearranged into a new set of vectors

$$
Q_{i}^{8}=\left\|\left\{q_{i}^{\theta}(y)\right\}_{Y_{0} I}\right\|^{N_{y^{0}} I}
$$

Since all components of $Q_{i}^{0}$ are calculated from statistically independent sets of data, the vector has a diagonal covariance matrix. In fact we have
where $\left\{F_{y}^{-1}\right\}_{i_{0} i}$ is the $i^{\text {th }}$ diagonal element of the inverse matrix $\frac{\mathrm{F}_{3}^{-1}}{}$.

We define a matrix of weights

$$
\underline{X}_{i}=G_{Q_{i}^{-q}}^{-1}
$$

and a vector of estimates for coincidence coefficients.

$$
A_{i}^{i 8}=\left\|\left\{a_{i j}^{n}\right\}_{j, I}\right\|^{n_{y, I}}
$$

Then by the use of model equation (3.3.10) we can obtain the soluteions

$$
A_{i}^{n}=K_{i}^{-1} \underline{Y}^{T} \underline{U}_{i} Q_{i}^{0}
$$

where the positive definite matrix $\mathrm{K}_{\mathrm{i}}$ is defined by

$$
\underline{K}_{i}=\underline{Y}^{T} \underline{W}_{i} \underline{Y} .
$$

As before, we have the result that the solutions are unbiased estimates of the true coefficients $a_{i j}$ and that the covariance matrices are given by

$$
\stackrel{C}{A}_{i}^{n}=\frac{K}{K}_{i}^{-1}
$$

From this last expression one can obtain all variances of the estirnoates $a_{i j}^{\prime \prime}$ and also the covariances between terms with equal i. The other covariances are generally not zero and are as yet undetermined. In Appendix $V$ it is shown that the full covariance matrix of the vector

$$
A_{2}^{\prime \prime}=\left\|\left\{\left\|\left\{a_{i j}^{\prime \prime}\right\}_{j, I}\right\|^{A_{y}, 1}\right\}_{i_{9} I}\right\| \|^{n_{x^{2}} 1} \quad 3.3 .38
$$

is given by

$$
{\underset{A}{A}}_{A_{2}^{1 \prime}}=\left\|\left\{\underline{K}_{i}^{-1} \underline{Y}^{T} \underline{U}_{i j} \underline{Y} \mathbb{K}_{j}^{-1}\right\}_{i, j}\right\|
$$

The diagonal matrix $\mathbb{U}_{i j}$ is defined by

$$
\underline{U}_{i j}=\left\|\left\{p_{i j}(y) \delta_{k y}\right\}_{k, y}\right\|^{N_{y}, N_{y}}, \quad, \quad 3.3 .40
$$

where
$\rho_{i j}(y)=$

$$
\frac{\left\{F_{\mathrm{y}}^{-1}\right\}_{i, j}}{\left\{F_{\mathrm{y}}^{-1}\right\}_{i, i} \quad\left\{F_{y}^{-1}\right\}_{j, j}}
$$

Consider now the diagonal submatrices of ${\underset{A}{A I \prime}}^{2}$. Putting $j=i$ we obtain

$$
\begin{aligned}
\left\{\underline{K}_{i}^{-1} \underline{Y}^{T} \underline{U}_{i i} \underline{Y} \underline{K}_{i}^{-1}\right\}_{i, i} & =\underline{K}_{i}^{-1} \underline{Y}^{T} \underline{W}_{i} \underline{Y} \underline{K}_{i}^{-1} \\
& =\frac{K_{i}^{-1}}{i}
\end{aligned}
$$

This result agrees with the covariance matrices given by equation (3.3.37). The off-diagonal sumatrices of ${\underset{-A}{-10}}^{\text {grive covariances which }}$ were not obtainable before, namely covariances between coefficients $a_{i j}^{\prime \prime}$ having different values of $i$.

According to the inequality (3.2.12), the variances in $A_{2}^{11}$ are greater than or equal to the corresponding variances in the full solution vector $A_{2}{ }^{\circ}$. An exact comparison of variances can be made only after inverting the full solution design matrix $\underline{B}_{2}$. This, however, may be irapractical in a routine calculation since one of the reasons for going to the solution by parts is precisely to avoid such a matrix inversion. It is thereiore desirable to have some simple method of determining lower limits to the variances in $A_{2}^{0}$ The design matrix $\underline{B}_{2}$ can be written in terms of submatrices $\underline{B}_{i j}$.

$$
\underline{B}_{2}=\left\|\left\{\underline{B}_{i j}\right\}_{i, j}\right\|^{n_{x^{\prime}} x}
$$

$B_{i j}=\left\|\left\{x_{0} \Sigma_{0} X_{i}(x) X_{j}(x) Y_{p}(y) Y_{q}(y) / M_{2}(x, y)\right\}_{p, q}\right\| \begin{aligned} & y^{n}, n_{y} \\ & 0.3 .44\end{aligned}$ It can be shown that for variances $D\left(a_{1, \ell}^{\prime}\right)$ in the vector $A_{2}^{\prime}$ the following inequalities hold (see Appendix V)

$$
D\left(a_{k \ell}^{0}\right) \geqslant\left\{\underline{B}_{k k}^{-1}\right\}_{\ell, \ell} \geqslant \frac{1}{\left\{\underline{B}_{2}\right\}_{i, i}}
$$

where the index $i$ is given by

$$
i=(k-1) n_{y}+2
$$

In expression (3.3.45) the second term represents the $l^{\text {th }}$ diagonal element in the inverse of $k^{\text {th }}$ diagonal submatrix ${\underset{W}{k} k^{*}}$ It is the
stronger of the two given lower limits. The weaker limit (last term in (3.3.45))is just the reciprocal of $i^{\text {th }}$ diagonal element in matrix $\mathrm{B}_{2}{ }^{\circ}$

These limits can be used to establish upper bounds on the magnification of variances introduced through the solution by parts. Some examples of actual variance magnifications and calculated upper limits are given below.

The models chosen for the purpose of comparing variances are based on the simple scheme of two r-rays in cascade. A computer program for generating response functions corresponding to a $3^{\prime \prime} \times 3^{\prime \prime} \mathrm{NaI}(\mathrm{Tl})$ detector at 10 cm . from a $\gamma$-ray source was used (Archer ${ }^{(47)}$ ). Response functions were assumed to be identical in both dimensions and were generated for an assumed analyzer gain of $40 \mathrm{keV} /$ channel. Assuming that $a_{21}=a_{22}=0$ and that $a_{12}=a_{21}=100$, the model coincidence spectrum is

$$
M_{2}(x, y)=100\left[X_{1}(x) x_{2}(y)+X_{2}(x) x_{1}(y)\right] \cdot 3.3 .46
$$

A set of 13 models characterized by different response functions $X_{2}(x)$ was used. Function $X_{1}(x)$ was kept fixed with its photo-peak centeredin channel 20. The photo-peak position of $X_{2}(x)$ was varied from channel 30 to 20.1 in steps'as shown in column 1 of Table III.

The spectrum field size used was $33 \times 33$ channels. With the given response functions (having shapes similar to the ones in Fig. ure 13) the spectrum has some regions of zero content. Consequently the weighting function $I / M_{2}(x, y)$ contains singularities. This difficulty was eliminated'by slightly modifying the response functions with the addition of a constant equal to $10^{-3}$.

TABLE III

Comparison of Variances

| Peak Positions, Channel Nos. | Singularity <br> Parameter <br> s | $\begin{aligned} & \text { Average } \\ & \text { Variance } \\ & D\left(a_{i j}^{b}\right) \end{aligned}$ | Average Variance Ratio $\quad \mathrm{r}$ | Stronger Upper Limit $\ell_{s}$ | Weaker <br> Upper <br> Limit <br> $\ell_{w}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 20,30 | . 465 | 130 | 1.030 | 1. 227 | 1.502 |
| 20,29 | . 4.45 | 136 | 1.033 | 1.240 | 2.538 |
| 20,28 | . 424 | 143 | 1.036 | 1. 255 | 1.582 |
| 20,27 | . 407 | 148 | 1.040 | 2.271 | 2.620 |
| 20,26 | . 421 | 144 | 1.041 | I. 262 | 1.601 |
| 20,25 | . 495 | 123 | 2.042 | 1. 218 | 1.466 |
| 20,24 | . 590 | 103 | 1.040 | 1.167 | 1.339 |
| 20,23 | . 592 | 103 | 1.042 | 1. 163 | 1.336 |
| 20,22 | . 379 | 154 | 2.054 | 1. 295 | 1.661 |
| 20,21 | $3.71 \times 10^{-2}$ | 702 | 1.036 | 2.19 | 5.13 |
| 20, 20.5 | $4.41 \times 10^{-4}$ | $8.05 \times 10^{3}$ | 2.004 | 6.68 | 47.2 |
| 20,20.2 | $2.25 \times 10^{-6}$ | 1. $23 \times 10^{5}$ | 1.0004 | 26.2 | 667 |
| 20,20.1 | $2.2 \times 10^{-7}$ | $4.16 \times 10^{5}$ | 1.0001 | 46.9 | 2240 |

Variances of the full solution estimates $a_{i j}^{\prime}$ and the solution by parts estimates $a_{i j}^{n}$ were calculated for each model. Column 3 of Table III gives the average variance

$$
\overline{D\left(a_{i j}^{0}\right)}=\frac{1}{4} \sum_{i, j=1}^{2} D\left(a_{i j}^{q}\right)
$$

which is seen to increase quite rapidly as the two photo-peaks reach close proximity. The average ratio of variances

$$
r=\frac{1}{4} \sum_{i, j=1}^{2}\left[D\left(a_{i j}^{\prime \prime}\right) / D\left(a_{i j}^{\prime}\right)\right]
$$

is given in column 4. It is seen that this ratio is very close to unity, the greatest departure fron that value being only 5.4\%.

The second column of Table III provides a measure of the approach to singularity of design matrix $\underline{B}_{2}$. Since $\underline{B}_{2}$ is positive definite and symmetric, the ratio of its determinant to the product of diagonal elements cannot be greater than unity. In equation form we have the singularity parameter

$$
s=\frac{\operatorname{det}\left(B_{2}\right)}{\operatorname{Ti}_{i}\left\{\underline{B}_{2}\right\}_{i_{8} i}} \leq 1
$$

If $\underline{B}_{2}$ were diagonal this ratio would equal unity. However, diagonality cannot be had in $\underline{B}_{2}$ since the response fiunctions used do not possess any orthogonality properties. For photo-peak separations greater than one channel the value of $s$ is near 0.5 . As the separation is reduced, matrix $\underline{B}_{2}$ approaches singularity and the ratio $s$ tends to zero. Uncertainties in least squares solution become very large but the ratio of variances, $r$ appears to approach unity.

Upper limits for the ratio $r$ were calculated by the use of inequalities (3.3.45). Column 5 of Table III gives values of the stronger upper limit

$$
\ell_{s}=\frac{1}{4} \sum_{i, j=1}^{2}\left[D\left(a_{i j}^{n \prime}\right) /\left\{\underline{B}_{i i}^{-1}\right\}_{j, j}\right] \cdot 3.3 .50
$$

The weaker limit, given in column 6, is

$$
\ell_{w}=\sum_{i}^{\frac{1}{4}} \sum_{i, j=1}^{2} D\left(a_{j, j}^{\eta}\right) \sum_{x, y=1}^{33}\left[x_{i}^{2}(x) x_{j}^{2}(y) / M_{2}(x, y)\right] \cdot 3.3 .51
$$

For most of the models tested both linits remain close to unity and therefore are useful. However, these limits provide little information when s approaches zero. This limitation is not serious, since, in the cases where $s$ is very small, the uncertainties are so great that the practical application of a least squares solution may become unfeasible.

Variances resulting from the two methods of solution were calculated also for the example given below, which has four $\gamma$-ray transitions and consequently 16 coincidence coefficients $a_{i j}$. The average variance ratio $r$ was 1.01 . It would thus appear that in most practical cases the magnification of uncertainties introduced by the solution by parts is insignificant.

In order to illustrate the potential power of coincidence analysis an example of calculations was carried to conclusion. The hypothetical decay scheme used is shown in Figure 12. There are four $\gamma$-ray transitions indexed from 1 to 4 in order of their energies $410,600,650$ and 1250 keV . Intensities are indicated by the total numbers of transitions $\mathscr{N}_{i}$ which occurred during the time of experiment. The first transitio: $\gamma_{1}$ is shown to have occurr-

FIGURE le

ence of $10^{7}$ counts. Following this there are 2 branches. Transitions $\gamma_{2}$ and $\gamma_{3}$ in cascade have an occurrence of $70 \%$, whereas the single $\gamma_{4}$ takes up the remaining $30 \%$.

In order of increasing energy the total efficiencies used were 0.294, 0.293, 0.293, 0.283 for detector $x$ and $0.265,0.265$, 0.265, 0.254 for detector $y$. The parameter $k$ in equation (3.3.6) was fixed for $10 \%$ chance coincidences by setting $k \mathcal{N}_{1}=0.1$. Response functions were generated using the different gains of $20 \mathrm{keV} / \mathrm{ch}$. in the $x$-dimension and $25 \mathrm{keV} / \mathrm{ch}$. in the y -dimension. The set of functions for the $y$-dimension, $Y_{j}(y)$ is shown in Figure 13. A data field of 3100 channels was used with $N_{x}=62$ and $N_{y}=50$.

Model one-dimensional and two-dimensional (coincidence) spectra were calculated by the use of equations (3.3.1) through (3.3.6). Angular correlation effects were neglected by setting $\bar{W}_{i j}(\theta)=1$. To simulate experimental spectra statistical deviations sampled from appropriate normal distributions were added to the models. Resultant values were rounded off to the nearest integer to give the final channel content $M^{8}$. If $M^{\circ}$ came out negative, it was arbitrarily set to zero.

The one-dimensional spectram in the $y$-dimension, $M_{y}^{0}(y)$, is shown in Figure 14. All four response functions of Figure 13 show their presence in various intensities. A least squares calculation produced estimates $a_{y j}$ of $2.6528 \times 10^{6}, I_{0} 8537 \times 10^{6}, ~ I .8556 \times 10^{6}$ and $7.6189 \times 10^{5}$ in order of increasing $j$. The largest fractional deviation from model value $\varepsilon_{y j} \mathscr{N}_{j}$ occurred in the first coefficient $a_{y I}^{\prime}$ and was $0.11 \%$. The content of each channel in spectrum $M_{y}^{\prime \prime}(y)$ is quite large $\left(M^{\prime}>4000\right)$. Consequently the statistical


deviations which are of the order $\sqrt{M^{8}}$ represent a small percentage too small to be visible in Figure 14. With these "good statistics" the use of weighting factor $1 / x^{\prime}$ in the least squares calculation provided an adequate approximation to the true weights $1 / \mathrm{M}$. A sinnilar analysis of the spectrum $M_{X}^{\prime}(x)$ in the $x$-dimension yielded coefficients $a_{x i}^{\prime}$ of $2.9417 \times 10^{6}, 2.0473 \times 10^{6}, 2.0564 \times 10^{6}$ and $8.4843 \times 10^{5}$. The largest fractional deviation from the model value was $0.26 \%$ in the third coefficient. These coefficients will be of use in determining the absolute intensities $\mathscr{N}_{i}$ later on。

The generated coincidence spectrum $M_{2}^{\prime}(x, y)$ contains areas of very low channel content. For these channels $1 / M^{\circ}$ can be a very poor approximation of the true weight. An iterative process was used to obtain better weight estimates. Weights $1 / M^{8}$ were used in a preliminary least squares solution by parts. Channels with $M^{\prime}<5$ were ignored in the calculation. The ensuing coincidence coefficients were then substituted in model equation (3.3.5) to obtain the improved weight estimates $I / M^{18}$. A second least squares solution by parts (rejecting channels with $M^{98}<5$ ) yielded the final estimates $a_{i j}$ ". Comparison of results obtained with various weight estimates is made in the following section. Intermediate coefficients $q_{i}^{\prime}(y)$ are plotted in Figure 15. Each shown function corresponds to a spectrum in the $y$-dimension which is in coincidence with the $i^{\text {th }}$ transition recorded in detector $x$. Except for small contribution due to chance coincidences, we see only
$94$

transitions $\gamma_{2}: \gamma_{3}$ and $\gamma_{4}$ in $q_{1}^{\prime \prime}(y) ; \gamma_{1}$ and $\gamma_{3}$ in $q_{2}^{8}(y) ; \gamma_{1}$ and $\gamma_{2}$ in $q_{3}^{\prime}(y)$ and finally - only $\gamma_{1}$ in $c_{4}^{\prime}(y)$. These spectra are simpler than the one-dimensional spectrum of Figure 14 and therefore can be advantageous in stuàying response function shapes or in searching for new transitions not included in the old model.

Using equations (3.3.3), (3.3.4) and (3.3.6) we can obtain

$$
\frac{a_{i, j}}{a_{x i} a_{y j}}=\frac{\mathscr{N}_{j, j}}{\mathscr{N}_{i} \mathscr{N}_{j}}+k
$$

Since $\mathscr{N}_{\text {ii }}=0$, the diagonal elements in the above array determine the parameter k. This can be subtracted from the whole array leaving the simplified elements $\mathscr{N}_{i j} /\left(\mathscr{N}_{i} \mathscr{N}_{j}\right)$. The 4 by 4 matrix of these simplified elements (calculated from the results of least squares solutions) is given in Table IV. The value of the chance parameter $k$ was calculated to be $k^{8}=1.017 \times 10^{-8}$ 。 which is $1.7 \%$ above the model value.

Diagonal elements of Table IV are expected to be zero. The degree of their departure from this value is a rough indication of the overall accuracy of the entries in the table. The large values of second, third and fourth elements in row $I$ and column 1 indicate that transition $\gamma_{1}$ has cascade coincidences with all other transitions. Thus $\gamma_{1}$ cannot have any competing branches. Further examination of Table IV reveals that transitions $\gamma_{2}$ and $\gamma_{3}$ are in cascade, whereas $\gamma_{4}$ forms a separate branch. These conclusions are in agreement with the originally assumed decay scheme of Figure 12.

$$
\text { For } \mathscr{N}_{i} \geqslant \mathscr{N}_{j} \text { it is possible to write }
$$

## TABIE IV

Matrix of Values $\left[\mathscr{N}_{i j}^{\prime} /\left(\mathcal{N}_{i}^{\prime} \mathcal{N}_{j}^{\prime}\right)\right] \times 10^{8}$

|  | $j$ | 1 | 2 | 3 |
| :---: | :---: | :---: | :---: | :---: |
| $i$ | 1 | -0.018 | 10.023 | 9.932 |
| 2 | 9.982 | 0.018 | 14.276 | -0.039 |
| 3 | 9.923 | 14.223 | -0.011 | -0.033 |
| 4 | 10.032 | -0.003 | -0.041 | 0.010 |

$$
\begin{align*}
\mathcal{N}_{i j} /\left(\mathcal{N}_{i} \mathcal{N}_{j}\right) & =\mathscr{N}_{i} \alpha_{i j} /\left(\mathscr{N}_{i} \alpha_{i j} \mathcal{N}_{i}\right) \\
& =1 / \mathcal{N}_{i}
\end{align*}
$$

where $\alpha_{i j}$ is the fraction of transitions $\gamma_{i}$ leading to transitions $\gamma_{j}$. Averaging elements with ( $i, j$ ) equal to ( 1,2 ) $(1,3),(1,4),(2,1)$, $(3,1)$ and $(4,1)$ we thus obtain the value $9.978 \times 10^{-8}$, the reciprocal of which yields

$$
\mathcal{N}_{I}^{0}=1.0022 \times 10^{7}
$$

This result is only $0.22 \%$ above the model value $\mathcal{N}_{1}$. Similarly, from elements $(2,3)$ and $(3,2)$ we obtain the estimate

$$
\mathscr{N}_{2}^{0}=\mathscr{N}_{3}^{0}=7.018 \times 10^{6}
$$

which exceeds the model value $7 \times 10^{6}$ by only $0.26 \%$. An estimate of the intensity in the competing branch $\gamma_{4}$ is obtained by simply taking the difference $\mathcal{N}_{1}^{\prime}-\mathcal{N}_{2}^{8}$.

These conclusions, derived from Table $I V$, required no knowledge of the detector efficiencies. Since the latter quantities are either unknown or difficult to obtain, their elimination can be a great advantage in studies of nuclear decay schemes.

### 3.4 Choice of Weighting Function

The general least squares solution discussed in the first two sections of the present chapter requires advance knowledge of the weighting function. This information is often not available and one is forced to resort to the use of some estimate of the proper weights. When the data vector is normally distributed, the required weights are equal to the reciprocals of variances. Use of these weights accounts for the fact that various data points have un-
equal uncertainties and hence must be assigned different degrees of irportance in the least squares analysis.

In nuclear counting experiments the content $M^{\prime}$ in any given channel has a Poisson frequency function (see Chapter I). If the mean, or expectation, value is denoted by $M$, then the variance is given by $E\left(M^{\prime}-M\right)^{2}=M$. For large values of $M$ (say $M>5$ ) the Poisson frequency function assumes a nearly Gaussian shape; consequently the maximum likelihood criterion leads to the method of least squares.

Since the true weights $1 / M$ are unknown, a common practice is to use the estimate $1 / M^{\prime}$. This can lead to some difficulties, especially when the spectrum contains channels with low counts. The value of $M^{8}$ may be zero, leading to a meaningless infinite weight even when the true weight 1/M is quite small. This necessitates special treatment (or complete exclusion) of channels with zero content. Moreover, the weight expectation $E\left(I / M^{\prime}\right)$ is generally different from the more realistic weight I/E(M'). Using the Poisson frequency function

$$
p\left(M^{\prime}, M\right)=\left(M^{M P} / M^{\prime}!\right) e^{-M}
$$

we have

$$
E\left(I / M^{n}\right)=\sum_{M^{\prime}=M_{\&}}^{\infty} p\left(M^{\prime}, M\right)\left(I / M^{\prime}\right)
$$

where $\mathrm{M}_{2}$ is the lowest channel content not rejected from the least squares fit. Ratios of $E\left(I / M^{\prime}\right)$ to the true weight $I / M$ are shown by the solid curves in Figure 16. The six curves were calculated for $M_{d}=1,2,3,4,5$ and 10. It is seen that for most values of $M$ the estimates $1 / M^{r}$ tend to be too high. For $M_{2}=1$ the excess is almost $30 \%$ at $M=4$ and gradually diminishes with increasing M. At the high value $M=100$ the excess is about $1 \%$.


It appears that a better estimate of weights can be obtained by using $I /\left(M^{\prime}+1\right)$. With this estimate channels with zero content require no special consideration. For a given value of $M$ we can express the weight expectation

$$
\begin{align*}
E\left[I /\left(M^{r}+1\right)\right] & =\sum_{M^{p}=M^{2}}^{\infty} p\left(M^{p}, M\right)\left[1 /\left(M^{p}+1\right)\right] \\
& =\frac{1}{M}\left[1-e^{-M} \sum_{n^{-}=0}^{M_{2}} \frac{M^{n}}{n!}\right] .
\end{align*}
$$

When no channels are rejected (i.e. $M_{l}=0$ ) the ratio of $\left.E\left[1 / M^{0}+1\right)\right]$ to. $I / M$ is simply $I-e^{-M}$, which rapidly approaches unity with increasing $M_{0}$ These ratios are shown in Figure 16 by the dotted curves calculated for $M_{2}=0, I, 2,3,4,5$ and 10. It is apparent that the averages $E\left[I /\left(M^{9}+1\right)\right]$ approach true weights $I / M$ much faster than is done by values $E\left(I / M^{P}\right)$ of the solid curves.

Better weights than either of the two above estimates can be obtained by a process of iteration. The above weights can be used in an initial solution for the least squares parameters. The latter can then be substituted in the model equation to get model spectrum estimates $M^{\prime \prime}$ for each point in the data field. A second least squares calculation can then be performed using the more accurate weight estimates $1 / M^{19}$.

Various weighting schemes were tested by using the previously discussed (section 3.3) two-dimensional spectrum with 16 coefficients $a_{i j}{ }^{\circ}$ Least squares solutions by parts were performed with three varieties of weights and a number of acceptance limits $M_{\mathscr{C}^{\bullet}}$ A basis of comparison was obtained by defining the dispersion parameter

$$
a=\sum_{i, j=1}^{4} \frac{\left(a_{i j j}^{n}-a_{i, j}\right)^{2}}{\sigma_{i j}^{2}}
$$

The same set of variances $\sigma_{i j}^{2}$ was used in all calculations of $d$. This set was obtained from the covariance matrices $\mathrm{K}_{\mathrm{i}}^{-1}$ in the solution with weights $W^{\prime}=1 / M^{\prime \prime}$ and acceptance $\operatorname{limit} M_{\&}=5$. In a good fit $\alpha$ is expected to have a value near 16. Larger values will result when the estimates $a_{i j}^{\prime \prime}$ have excessive deviations from the model values $a_{i j}$ 。 Results are presented in Table Ve Column $I$ shows the acceptance limit $M_{2}$ ranging from 0 to 25. Columns 2 and 3 give the number of channels which fell below the limit $M_{\&}$ and were excluded from the least squares calculation. Results in column 2 obtain in calculations with weights $W^{\prime}=I / M^{\prime}$ and $W^{\prime}=I /\left(M^{\prime}+1\right)$, whereas column 3 corresponds to the weights $W^{\prime}=I / M^{\prime \prime}$. The last three columns give the parameter $\alpha$ obtained with various weights. The iterative weights $1 / M^{\prime \prime}$ were calculated from the results of an initial solution with $W^{\prime \prime}=I / M^{\prime \prime}$ and $M_{k}=5$.

In all cases the weights $I /\left(M^{9}+1\right)$ appear superior to the weights $I / M^{\prime}$ in that their associated dispersion parameter $d$ has smaller values. As expected, the difference diminishes with increasing $M_{\ell^{\circ}}$ Also expected is the observed general tendency for $\alpha$ to decrease at higher values of $M_{\mathcal{L}^{\prime}}$. On the other hand, the iterative weights $I / M^{\prime \prime}$ yield values of $d$ relatively independent of $M_{2}$ within the range of values used. It thus appears that in cases where channels with very low content must be used, the iterative weights $1 / M^{\prime \prime \prime}$ are preferable. When $M_{\&}$ can be set at 15 or 20 , the weights $I /\left(M^{\prime}+1\right)$ seem to be adequate. The use of weights $I / M^{\prime}$ is not recommended for any condition.

## TABLE $V$

Test of Various Weighting Functions

| Acceptance <br> Limit <br> $M_{2}$ | No. of Channels <br> Rejected out of 3100 |  | Dispersion Parameter d |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\mathrm{M}^{8}<\mathrm{M}_{\ell}$ | $M^{\prime \prime}<M_{l}$ | $W^{8}=I / M^{9}$ | $W^{8}=1 /\left(M^{8}+1\right)$ | $W^{\prime}=1 / M^{\prime \prime}$ |
| 0 | 0 | - | - | 55.0 | - |
| 1 | 167 | 193 | 38.5 | 26.8 | 12.5 |
| 5 | 588 | 570 | 25.3 | 24.9 | 13.6 |
| 10 | 750 | 765 | 33.4 | 32.3 | 14.7 |
| 15 | 836 | 835 | 24.9 | 14.7 | 14.9 |
| 20 | 890 | 897 | 16.0 | 15.9 | 14.7 |
| 25 | 939 | 947 | 13.8 | 13.6 | 14.7 |

### 3.5 Application to Nonlinear Models

The models considered thus far had the property of being linear functions of the least squares parameters. This led to linear normal equations which could be solved by relatively simple matrix methods. However, the use of more complicated model functions usually results in a set of normal equations which are very difficult to solve. Some modification in approach is required to facilitate the practical solution for the pertinent parameters.

Nonlinear models are essential for accurate determination of energies in nuclear spectra. Many detector response functions involve a peak which has a shape well approximated by a Gaussian function (e.g. the photo-peak in Figure 1). In order to determine the central position-and hence the energy associated with the Gaussian peak - it is necessary to perform a least squares calculation based on this nonlinear model.

One practical method of solution $(40,44,47)$ Iinearizes the model function by expanding it in a Taylor series and neglecting terms which are higher than the first order. With this method it is necessary to provide initial estimates of the least squares parameters. If the estimates are good, then the required corrections turn out to be small and the linear approximation is fairly accurate. The calculations can be repeated with the new corrected values being used as estimates. After several such interations the calculated correctionsusually become insignificant and the process can be stopped.

We define a general model function

$$
M(x)=M\left(a_{1}, a_{2}, \cdots, a_{n} ; x\right)
$$

and denote by $M^{\circ}(x)$ the function obtained when initial estimates $a_{i}^{0}$ are substituted in (3.5.1). The function $M(x)$ can be approximated by the truncated Taylor series

$$
M^{1}(x)=M^{0}(x)+\sum_{i=1}^{n} \frac{\partial M^{0}(x)}{\partial a_{i}^{0}} \Delta^{I} a_{i}
$$

where

$$
\Delta^{l} a_{i}=a_{i}-a_{i}^{0}
$$

We can now state the least squares condition

$$
\sum_{x=1}^{N} \frac{1}{M^{1}(x)}\left[M^{P}(x)-M^{0}(x)-\sum_{i=1}^{n} \frac{\partial M^{0}(x)}{\partial a_{i}^{0}} \Delta^{I} a_{i}\right]^{2} \rightarrow \min _{3} 3.5 .4
$$

where, as before, the data vector is represented by $M^{\text { }}(x)$. Treating the corrections $\Delta^{l} a_{i}$ as our least squares parameters, we obtain the set of $n$ normal equations
$\sum_{x=1}^{N} \frac{1}{M^{1}(x)}\left[M^{0}(x)-M^{0}(x)\right] \frac{\partial M^{0}(x)}{\partial a_{k}^{0}}=\sum_{x=1}^{N} \frac{I}{M^{I}(x)} \frac{\partial M^{0}(x)}{\partial a_{k}^{0}} \sum_{i=1}^{n} \frac{\partial M^{0}(x)}{\partial a_{i}^{0}} \Delta^{I^{\prime}} a_{i}$,
where the index $k$ can have values $L_{,} 2, \ldots, n_{0}$
These results can be gathered into a neater matrix representation. Using the symbolism of section 3.4 we define the matrices and vectors

$$
\begin{aligned}
& \underline{M}^{I}=\left\|\left\{M^{I}(x)-M^{0}(x)\right\}_{X_{9} I}\right\|^{N_{8} I} \\
& \underline{D}^{0}=\left\|\left\{\frac{0}{2.5} \frac{0.6}{\partial a_{j}^{0}}\right\}_{x_{9} i}\right\|^{N_{8} n}
\end{aligned}
$$

$$
\begin{aligned}
& \Delta^{1}=\left\|\left\{\Delta^{1} a_{i}\right\}_{i_{2} I}\right\| \begin{array}{lll}
n_{0} 1 & 3.5 .8
\end{array} \\
& W^{I}=\left\|\left\{\frac{\delta_{i x}}{M^{2}(x)}\right\}_{i_{0} x}\right\|^{N_{8} N}, \quad 3.5 .9 \\
& \underline{B}^{\circ}=\underline{D}^{D^{T}} \underline{N}^{T} \underline{D}^{\circ} \\
& =\left\|\left\{\begin{array}{llll}
\sum_{x=1}^{N} & \frac{I}{M^{I}(x)} & \frac{\partial M^{0}(x)}{\partial a_{j}^{0}} & \frac{\partial M^{0}(x)}{\partial a^{0}{ }_{j}}
\end{array}\right\}_{i_{8} j}\right\|^{n \cdot n} \quad 3.5 .10
\end{aligned}
$$

The model equation (3.5.2) becomes now

$$
\underline{M}^{I}=\underline{D}^{0} \underline{\Delta}^{I}
$$

and the normal equations (3.5.5) can be replaced by

$$
\underline{B}^{0} \underline{\Delta}^{I^{\prime}}=\underline{D}^{0^{T}} \underline{W}^{I} \underline{M}^{I^{\prime}}
$$

where

$$
\underline{M}^{I^{\prime}}=\left\|\left\{M^{\prime}(x)-M^{0}(x)\right\}_{X_{0} 1}\right\|^{N_{2} I}
$$

Solution for the vector of corrections $\Delta^{I^{\prime}}$ can be effected by the inversion of design matrix $\mathbb{B}^{0}$. The components of this vector are added to the initial estimates $a_{i}^{0}$ to obtain a new set of estimates

$$
a_{i}^{I^{\prime}}=a_{i}^{0}+\Delta^{I_{a}^{\prime}}
$$

If all superscripts are increased by unity, then the same set of equations (3.5.2) to (3.5.14) will describe the second iteration. The iterations are continued until the added corrections in equation
(3.5.14) have negligible effect.

In general there is no assurance that the process converges to give estimates $a_{i}^{\prime}$ which will minimize the function

$$
\therefore 2=\sum a^{2}+\quad\left[\left(u^{\prime}\right)-M\left(a_{1}^{0}, a_{2}^{0}, \ldots, a_{n}^{0} ; x\right)\right]^{2} \cdot 3.5 .15
$$

There may exist local minima which provide non-optimum estimates $a_{i}^{\prime}$, or the process may not converge at 2il. Much depends on the model function used and the choice of initial estimates. Fortunately a Gaussian model has good convergence properties and its parameters are usually easy to estimate by a visual inspection of the spectrum. Thus the method is well suited for determination of transition energies. However, when the model includes more than one peak located within the width $\sigma$, the results are not aiways predictable. In such cases use might be made of the "parabolic method" discussed in detail by Archer ${ }^{(47)}$. Near its minimum value $\chi^{\prime 2}$ is nearly a quadratic function of the parameters $a_{i}$. By calculating $\chi^{8}$ at various values of $a_{i}^{0}$ and fitting the results to quadratic functions it is often possible to find a set of parameters $a_{i}^{\prime}$ which will minimize $\chi^{{ }^{2}}$.

In general the estimates $a_{i}^{0}$ are not normally distributed when the model function is nonlinear. Their exact distributions are not easily determined and may require numerical Monte Carlo calculations. However, if certain conditions are satisfied, the actual distributions can be nearly normal. Consider a Taylor series expansion of $M(x)$ about the unknown true values of parameters
$a_{i}$. We have

$$
\begin{align*}
M\left(a_{i}^{\prime}, a_{2}^{\prime}, \ldots a_{i n}^{\prime} ; x\right) & =M(x)+\sum_{i=1}^{n} \frac{\partial M(x)}{\partial a_{i}}\left(a_{i}^{\prime}-a_{i}\right) \\
& + \text { (higher order terms }) .
\end{align*}
$$

If we assume now that the variances in estimates $a_{i}$ are small, then over the range of most probable values of $a_{i}^{\prime}$ the differences ( $a_{i}^{\prime}-a_{i}$ ) will also be small. Consequently we can neglect the higher order terms and expression (3.5.16) becomes linear in the parameter estimates $a_{i}^{\prime \prime}$. When the initial estimates $a_{i}^{0}$ are not too different from the true values, the iterative solution is independent of the values $a_{i}^{0}$ and converges to the same set $a_{i}^{\prime}$. Therefore we can assume that our initial estimates were the true values $a_{i}$ without altering the distribution in $a_{i}$. With the Iinear approximation in (3.5.16) only one iteration will be required to obtain the least squares estimates $a_{i}^{\prime}$ and this solution will have the same statistical properties as the solutions of the linear model in sections 3.1 and 3.2. Consequently, we can say that the estimates $a_{j}$ are normally distributed, that they have expectation values given by

$$
E\left(a_{1}^{\prime}\right)=a_{1}
$$

and that their covariance matrix is

$$
\underline{G}_{A^{\prime}}=\underline{B}^{-1}
$$

where

$$
\underline{B}=\left\|\left\{\sum_{x=1}^{N} \frac{1}{M(x)} \frac{\partial M(x)}{\partial a_{i}} \frac{\partial M(x)}{\partial a_{j}}\right\}_{i, j}\right\|^{n, n} \cdot 3.5 .19
$$

Since Gaussian functions find frequent application in non-
linear least squares calculations, it is perhaps worthwhile to derive an explicit expression for the covariance matrix of this particular model. As our model function we take the expression

$$
M(x)=a_{1} \exp \left[-\frac{\left(x-a_{2}\right)^{2}}{2 a_{3}^{2}}\right]
$$

and assume that the width $a_{3}$ is sufficiently large to permit. replacement of the summation in (3.5.19) by an integral sign.

We shall consider two particular cases. In the first instance it will be assumed that the available data extends over the full range of the Gaussian peak and the limits of integration used will be $-\infty$ to $+\infty$. Another case of interest occurs when only the right half of the peak approximates a Gaussian shape. For example, the response shapes of $\mathrm{NaI}(T \mathrm{l})$ detectors may contain small angle scattering events and contributions from the Compton scattering distribution which tend to distort the left half of the Gaussian photo-peak. For this case only the right half of the peak will be considered by using the limits of integration $a_{2}$ to $\infty$

We start by considering the full Gaussian peak. Substitution of $M(x)$ and its partial derivatives in expression (3.5.19) yields the design matrix

$$
\underline{B}=\left\|\begin{array}{ccc}
\frac{N_{0}}{a_{1}^{2}} & 0 & \frac{N_{0}}{a_{1} a_{3}} \\
0 & \frac{N_{0}}{a_{3}^{2}} & 0 \\
\frac{N_{0}}{a_{1} a_{3}} & 0 & \frac{3 N_{0}}{a_{3}^{2}}
\end{array}\right\|,
$$

where $N_{0}$ is the total number of counts under the peak, given by

$$
N_{0}=\sqrt{2 \pi} \quad a_{1} a_{3}
$$

The matrix is easily inverted to obtain the covariance matrix

$$
{\underline{c_{A}}}^{\prime}=\left\|\begin{array}{ccc}
\frac{3 a_{1}^{2}}{2 N_{0}} & 0 & \frac{a_{1} a_{3}}{2 N_{0}} \\
0 & a_{3}^{2} & 0 \\
\frac{a_{1} a_{3}}{2 N_{0}} & 0 & \frac{a_{3}^{2}}{2 N_{0}}
\end{array}\right\|
$$

Of some special interest is the central covariance matrix element $a_{3}{ }^{2} / N_{0}$ which provides the sariance in estimate of the peak position $a_{2}, \sigma_{a_{2}}^{2}$. To take a practical example, consider a Gaussian peak consisting of $N_{0}=10^{6}$ counts. Its full width at half maximum is given by $2.35 a_{3}$ and its central position is determined within an accuracy of $\pm a_{3} / \sqrt{N_{0}}$. The ratio of position uncertainty to the peak width is $1 / 2350$. This high precision demonstrates the power in the method of least squares analysis.

Next we turn to the case where usable data covers only half of the peak. Using the integration limits $a_{2}$ to $\infty$ we obtain the design matrix

$$
\underline{B}=\left\|\begin{array}{ccc}
\frac{N_{0}}{2 a_{1}^{2}} & 1 & \frac{N_{0}}{2 a_{1} a_{3}} \\
1 & \frac{N_{0}}{2 a_{3}^{2}} & \frac{2 a_{1}}{a_{3}} \\
\frac{N_{0}}{2 a_{1} a_{3}} & \frac{2 a_{1}}{a_{3}} & \frac{3 N_{0}}{2 a_{3}^{2}}
\end{array}\right\| .
$$

After a considerable amount of algebraic manipulation it is posso ible to obtain the expression for the inverse matrix.

$$
\underset{\sim}{C}=\left|\begin{array}{cc:c:cc}
\left(\frac{3 \pi-8}{\pi-3}\right){\frac{a_{1}}{2}}_{N_{0}}^{2} & -\frac{1}{(\pi-3)} & \frac{(4-\pi)}{(\pi-3)} & \frac{a_{1} a_{3}}{N_{0}} \\
\hdashline-\frac{1}{(\pi-3)} & \frac{2 \pi}{(\pi-3)} & \frac{a_{3}}{N_{0}} & \frac{-\sqrt{2 \pi}}{(\pi-3)} & \frac{a_{3}}{N_{0}} \\
\hdashline \frac{(4-\pi)}{(\pi-3)} & \frac{a_{1} a_{3}}{N_{0}} & \frac{-\sqrt{2 \pi}}{(\pi-3)} & \frac{a_{3}}{N_{0}} & \frac{(\pi-2)}{(\pi-3)} \\
\hdashline N_{0} & \frac{a_{3}}{N_{0}}
\end{array}\right|
$$

From the above result we see that the use of only half the peak has increased the variance $\sigma_{a}^{2}$ by a factor $2 \pi /(\pi-3)=44.3$. The uncertainty, or standard deviation, is increased hy the factor $\sqrt{44.3}=6.66$. Uncertainties in the estimates of $a_{1}$ and $a_{3}$ have increased by the factors 2.59 and 4.02 respectively.

In the foregoing discussion it was assumed that the Gaussian widith $a_{3}$ is sufficiently large to permit neglection of channel ine tegration effects. When $a_{3}$ is small it becomes necessary to replace equation (3.5.20) by the more accurate expression

$$
\begin{aligned}
M(x) & =a_{1} \int_{x-\frac{1}{2}}^{x+\frac{1}{2}} \exp \left[-\frac{\left(y-a_{2}\right)^{2}}{2 a_{3}^{2}}\right] d y \\
& =\frac{N_{0}}{2}\left\{\operatorname{erf}\left[\frac{x+\frac{1}{2}-a_{2}}{\sqrt{2} a_{3}}\right]-\operatorname{erf}\left[\frac{x-\frac{1}{2}-a_{2}}{\sqrt{2} a_{3}}\right]\right\}, 3.5 .26
\end{aligned}
$$

where the error function is defined by

$$
\operatorname{erf}(z)=\frac{2}{\sqrt{\pi}} \int_{0}^{z} e^{-t^{2}} d t
$$

It also becomes necessary to retain the sumation sign in design matrix definition (3.5.19), since the use of integrals would result
in a too crude approximationo Consequently simple expressions for the matrix elements are not available and the calculations were perm formed by numerical methods.

Elements of design matrix $\underline{B}$ were calculated on the assumption that data are available over the full range of the peak. Model eqation (3.5.26) contains the three constants $N_{0}, a_{2}$ and $a_{3}$ which were used as the least squares parameters. Upon inversion of the design matrix $B$ it was possible to obtain numerical values of $N_{0} \sigma_{a_{2}^{i}}^{2}$ independently of the value $\mathbb{N}_{0}$. Therefore it was necessary to vary only the values of $a_{2}$ and $a_{3}$ in order to study the behaviour of $\sigma_{a_{2}^{\prime}}^{2}$ at various parameter values.

The quantity $N_{0} \sigma_{a_{2}^{p}}^{2} / a_{3}^{2}$ is plotted as a function of width $a_{3}$ in Figure 17. Two curves are shown: one for the peak placed in the middle of a channel $\left(a_{2}=0\right)$ and one for a peak position halfway between channel centers ( $a_{2}=\frac{1}{2}$ ). Below the value $a_{3}=0.5$ the curves show considerable divergence. The curve with $a_{2}=\frac{1}{2}$ remains close to unity, conforming with the previous approximate res sult in covariance matrix (3.5.23), whereas the curve for $a_{2}=0$ displays a sharp rise with decreasing values of the width $a_{3}$.

The significance of the variance $\sigma_{a_{2}^{p}}^{2}$ at very low values of $a_{3}$ is questionable since the probability density function of estimate $a_{2}^{0}$ loses its Gaussian shape. In fact, as the width $a_{3}$ vanishes, the peak position becomes indeterminate within the limits of channel width and the probability density function of $a_{2}^{\prime}$ assumes the shape of channel profile $P(x)$, as defined by (2.5.28a). In this case the actual variance becomes


$$
\sigma_{a}^{2}=\int_{-\frac{1}{2}}^{2} \quad x^{2} d x=\frac{1}{12}
$$

and is independent of $N_{0}$. This channel variance manifests itself at all values of $a_{3}$ and the asymptotic expression for variance in estimate $a_{2}^{\prime}$ is

$$
\sigma_{2}^{2} \sim \frac{1}{N_{0}}\left[a_{3}^{2}+\frac{1}{12}\right]
$$

A curve corresponding to this asymptotic expression is included in Figure 17. It is seen that all three curves converge very rapidly at $a_{3}=0.5$.

In order to test the actual distributions of estimates $a_{2}^{\prime \prime}$ a number of Monte Carlo calculations were performed with various values of width parameter $a_{3}$. Spectra were generated by using values obtained from the model (3.5.26) and adding normally distributed random numbers having variances equal to $M(x)$. Three cases were studied with $a_{3}$ fixed at $4.0,0.5$ and 0.2 . In each case the model values of $N_{0}$ and $a_{2}$ remained constant at 100 counts and 0 chrnnels respectively. Figure 18 shows histograms of the least squares solutions $a_{2}^{\prime}$. Gaussian probability density functions with variances $\sigma_{a_{2}^{p}}^{2}$ taken from Figure 17 (curve with $a_{2}=0$ ) are also plotted for the sake of obtaining a comparison. Except for the case with the low value $a_{3}=0.2$, the estimates $a_{2}^{\prime}$ appear to be normally distributed.

Results of the above calculations can be summed up as follows. The nonlinear least squares solutions provide normally distributed estimates of Gaussian peak positions for values of $N_{0}$ as low as 100

FIGURE 18
HISTOGRAMS OF PEAK POSITION ESTIMATES $N_{0}=100$



counts (or lower) and values of width $a_{3}$ as low as 0.5 channelso When $a_{3}>0.5$, the variances in position estimates $a_{2}^{9}$ are very closely approximated by the asymptotic expression (3.5.29). Values of $a_{3}$ smaller than 0.5 channels should be avoided due to the possibility of a large loss in accuracy. Since $a_{3}$ is measured in channel widths, its value can be increased by redesigning the experiment to yield a finer channel mesh.

## CHAPIER IV

CONCLUSIONS

The methods of analysis presented in this thesis were discussed with special emphasis to application in nuclear spectroscopy. However, use of this emphasis was not intended to be indicative of a limitation in possible applications. The techniques presented can be generally applied to digitized statistical (or nonstatistical) data which contain the distorting effects of apparatus response functions. Statistical considerations were based on Poisson and Gaussian frequency functions which are found to occur in a wide class of physical measurements.

Two essentially different methods of approach were discussed. Both are concerned with the same problem of obtaining the physically meaningful parameters with least possible error. The matrix inversion approach requires no initial knowledge of the form the unfolded spectrum might take. It is thus convenient in cases where this information is unavailable or difficult to obtain. On the other hand, the least squares approach generally requires a model of the unfolded spectrum to the extent that the number of components and their approximate positions must be specified in advance. This often requires a detailed inspection of the original data followed by an educated guess of the pertinent unfolded spectrum parameters. However, the least squares approach has also some important advantages. Unlike the matrix model approach, it is not limited to integral values of the spectrum components. The removal of this restriction is important when accurate position determinations of narrow peaks
are required. In some applications the two approaches might be combined to take full advantage of their relative merits. A preliminary application of response matrix inversion may provide a useful visual representation of the unfolded spectrum and thus delineate the appropriate least squares model with a smaller element of uncertainty.

It was shown in Chapter II that the application of some inverse response matrices may lead to a gross magnification in statistical uncertainties. Two methods of combating this undesirable effect were suggested. It is possible to obtain a "partial inverse" matrix which will in effect replace the originally broad response functions, in the observed spectrum by a set of narrower ones of some specified shape. Since this "partial inversel" matrix has elements smaller in magnitude than the full inverse matrix, the statistical uncertainties in the unfolded spectrum can be greatly reduced. Another method of dealing with statistical devm iations is based on the a priori knowledge of non-negativity in spectral intensities. Corrections are added tothoobserved spectrum so as to render the unfolded spectrum positive (or zero) in every channel. The particular set of corrections to be added is determined by the condition that the sum of the weighted squares of their individual values should be minimized.

When the data vector is multiplied by the inverse matrix it is a simple matter to calculate the statistical variances in the resultant uafolded spectrum (see formula(2.6.1)). However, after application of non-negativity the determination of statistical uncertainties is somewhat more difficult. After non-negativity corrections are added to the observed spectrum, the result is a linear combination of response
functions with non-negative amplitude coefficients. These coefficients . one for each channel o represent the unfolded non-negative spectrum. The same unfolded spectrum would be obtained in a least squares calculation if one used a model equation which contains one response function with non-negative amplitude for each channel. The model could avoid negative amplitudes by the use of quadratic coefficients, in which case the nonlinear least squares techniques discussed in section 3.5 must be applied. Therefore, in principle, one could obtain the statistical properties of the completely unfolded non-negative spectrum from the inverse of the corresponding least squares design matrix.

Similas arguments can be applied to cases where a "partial inverse" matrix is used in conjunction with the application of non-negativity. In order to obtain a least squares analogy it is first necessary to construct a set of complementary response functions. These can be obtained by unfolding the residual response functions from the full width original response shapes. The corresponding least squares model is then constructo ed by placing one complementary response function with non-negative amplio tude in each data channel. Since the least squares solution gives a spectrum from which, in effect, the complementary functions have been unfolded, the remaining response shapes will be the same as in the application of partial matrix inversion. The covariance matrix (inverse of the design matrix) of this least squares model determines the statistical properties of the partially unfolded spectrum.

It is thus seen how the two basically different methods of approach can become identical under certain conditions. However, when the data fields are large, it is not practical to apply ordinary least squares
techniques to the solution of intensities in each channel. The required, inversion of the design matrix becomes unfeasible due to its large size. In these cases the response matrix approach can be of decisive advanto age when the inverse matrix elements are obtainable in closed form. The proper application of non-negativity conditions may require extensive numerical calculations based on methods of mathematical programming. However, it was demonstrated in section 2.6 that at least in some cases a very rapid iterative method can be used.

## APPENDIX I

## STATISTICAL PROPERTTES OF GHANNEL CONTENT

Assume that for a large number (approaching infinity) of identical experiments, the average count in channel $i$ is given by $m_{i}$. It is desired to find the relative probabilities of obtaining various numbers of counts $m_{i}$. To do this we postulate a large number $N\left(N \gg m_{i}\right.$ ) of "potential" counts. Each potential count has a probability $\frac{m_{i}}{N}$ of becoming a "real" count and is statistically independent of all others. The probability for a particular set of $m_{i}^{8}$ potential counts materializing and the remainder being not counted is given by

$$
\left(\frac{m_{i}}{N}\right)^{\frac{m_{j}}{2}}\left(1-\frac{m_{i}}{N}\right)^{N-m_{i}^{\prime}}
$$

Multiplying the last expression by the total number of such sets we obtain the probability of getting $m_{i}$ weal counts in channel i, namely

$$
\frac{N!}{\left(N-m_{i}^{\prime}\right)!\left(m_{j}^{\prime}\right)!}\left(\frac{m_{i}}{N}\right)^{\frac{m_{j}^{\prime}}{1}}\left(1-\frac{m_{j}}{N}\right)^{N-m_{i}^{\prime}} \quad . \quad \text { Al.I }
$$

We now let $N$ approach infinity, under which condition

asad

$$
\left(1-\frac{m_{i}}{N-m_{i}^{\prime}}\right)^{{ }^{N}} \rightarrow e^{-m_{i}}
$$

Therefore, in the limit, we obtain the required frequency function for $m^{\prime}{ }^{\prime}$,

$$
\begin{equation*}
p\left(m_{i}^{\prime}, m_{i}\right)=\frac{m_{i}^{\prime}}{\left(m_{i}^{\prime}\right)!} e^{-m_{i}} \tag{Al. 2}
\end{equation*}
$$

It should be noted that this frequency function also holds true for the contents of histogram bins, provided the histogram is made up of a large number of statistically independent experiments and is spread out over many bins of non-zero content. If the histogram is concentrated around one bin, then it is necessary to use expression (Al.1) with $N$ set equal to the total number of experiments.

When the average value $m_{i}$ is large, expression (Al.2) approaches a Gaussias irequency function. To show this we use Stirling's formula ${ }^{(54)}$

$$
\begin{equation*}
n!\sim e^{-n} n^{n} \quad \sqrt{2 \pi n} \tag{A1. 3}
\end{equation*}
$$

We can thus write

$$
\log p \sim\left(m_{i}^{i}+\frac{1}{2}\right) \log \left(\frac{m_{i}}{m_{i}^{j}}\right)-\left(m_{i}-m_{i}^{n}\right)-\frac{1}{2} \log \left(2 \pi m_{i}\right) .
$$

Next we use the series expansion (55)

$$
\log \left(\frac{m_{i}}{\frac{m_{i}^{g}}{j}}\right)=2\left\{\frac{m_{i}-m_{i}^{0}}{m_{i}+m_{i}^{0}}+\frac{1}{3}\left(\frac{m_{i}-m_{i}^{0}}{m_{i}+m_{i}^{g}}\right)^{3}+\ldots\right\}
$$

in which the cubic and higher order terms can be neglected since ${ }_{0}$ for values of $p$ significantly greater than zero, we have

$$
\frac{m_{i}-m_{i}^{1}}{m_{i}+m_{i}^{i}} \ll 1
$$

A. 4

We now can write the result

$$
\begin{aligned}
\log p & \sim\left(2 m_{i}^{\prime}+1\right) \frac{\left(m_{i}-m_{i}^{\prime}\right)}{\left(m_{i}+m_{i}^{\prime}\right)}-\left(m_{i}-m_{i}^{\prime}\right)-\frac{1}{2} \log \left(2 \pi m_{i}\right) \\
& =-\frac{\left(m_{i}^{\prime}-m_{i}\right)^{2}}{2 m_{i}} \frac{2 m_{i}}{\left(m_{i}+m_{i}^{\prime}\right)}-\frac{1}{2} \log \left(2 \pi m_{i}\right)
\end{aligned}
$$

When the strong inequality (Al.4) holds we have also the approximate relation

$$
\frac{2 m_{i}}{m_{i}+m_{i}^{\prime}} \approx 1
$$

in which case we can write the final asymptotic frequency function

$$
p\left(m_{i}^{\prime}, m_{i}\right) \sim \frac{1}{\sqrt{2 \pi m_{i}}} e^{\frac{-\left(m_{i}^{\prime}-m_{i}\right)^{2}}{2 m_{i}}}
$$

A comparison between the Gaussian and Poisson frequency functions is given in Figure 19 for various values of $m_{i}$. It is seen that the similarity is fairly good at $m_{i}=5$ and becomes increasingly better at higher values of $m_{i}$.


## APPENDIX II

CORRECTIONS FOR COINCIDENCE AND CHANCE SUMMING

The number of prompt coincidence summing events is given by

$$
N_{c s}={ }_{i, j}^{\Sigma_{j}} \varepsilon_{i} \quad \varepsilon_{j} \quad \mathcal{N}_{i j} \quad W_{i j},
$$

where $\mathscr{N}_{i j}$ is the number of transitions $\gamma_{i}$ followed by $\gamma_{j}, W_{i j}$ is the angular correlation function integrated over the solid angle $\Omega$ subm tended by the detector and $\mathcal{E}_{\mathrm{j}}$ is the overall detection efficiency for $\gamma_{i}$. Since efficiencies are proportional to $\Omega$, we can write

$$
\begin{equation*}
\varepsilon_{i}=k_{i} \Omega, \tag{A2. 2}
\end{equation*}
$$

which results in

$$
\begin{equation*}
N_{c s}=\Omega^{2} \sum_{i, j} k_{i} k_{j} \mathscr{N}_{i j} W_{i j} \tag{A2. 3}
\end{equation*}
$$

The number of single events is simply

$$
N_{s}=\Omega \underset{i}{\sum k_{i}} \mathcal{N}_{i} .
$$

Consequently the relative contribution from coincident sumaing, $N_{c s} / N_{s}$ is proportional to the solid angle $\Omega$. Therefore the effect can be reduced by placing the source at a greater distance from the detector.

Contribution due to random suming of uncorrelated events can be calculated from the following considerations. We have $N_{S}$ single events occurring during observation time T. Each of these events has the probability $2 \tau \mathbb{N}_{s} / T$ of being followed or preceded by another event within the electronic resolution time $\tau$. Therefore the expectation of the total random summing events is

$$
\begin{equation*}
N_{r s}=\frac{2 \tau}{T} N_{s}^{2} . \tag{A2. 5}
\end{equation*}
$$

The relative contribution $N_{r s} / N_{s}$ is thus proportional to the counting rate $N_{s} / T$ and the effect can be reduced by using weaker sources or smaller solid angle.

Practical considerations often require relatively high counting rates which could necessitate correction for summing effects. It is possible to make these corrections by using special experimental techniques or by applying appropriate post-experimental calculations.

Random sum pulses can be rejected to some extent by electronic circuitry which senses the pulse broadening ${ }^{(4)}$. However this method is not $100 \%$ efficient and has no effect on coincident sum events.

A powerful experimental technique can be applied by the use of two identical detectors ${ }^{(3)}$. The mixed output from both detectors is stored in one subgroup of the analyzer memory. Another memory subgroup receives the gated sum of pulses whenever both detectors respond in coincidence. The contents of the gated subgroup can then be subtracted from the mixed spectrum. A great advantage of this technique is that correction can be made for both random and coincident sum events ${ }^{(5)}$.

An expression for the spectrum of coincident sum events has been given by Heath ${ }^{(2)}$. Let the response functions to transitions $\gamma_{i}$ and $\gamma_{j}$ be given by $R\left(x, y_{i}\right)$ and $R\left(x, y_{j}\right)$ respectively. The two-dimensional probability density function for $\gamma_{i}$ to produce a pulse height $x$ and for $\gamma_{j}$ to produce a pulse-height $\mathrm{z}-\mathrm{x}$ is given by the product

$$
R\left(x, y_{i}\right) R\left(z-x, y_{j}\right) .
$$

The one-dimensional probability density function of obtaining two pulses
which add up to the value $z$ is given by the integral

$$
\int_{0}^{z} R\left(x, y_{i}\right) R\left(z-x, y_{j}\right) d x
$$

To obtain the total coincidence sum spectrum we must multiply this last expression by the number of coincidences $\mathscr{N}_{i j}$, the efficiencies $\varepsilon_{i}$, $\varepsilon_{j}$ and then sum over all values of $i$ and $j$. Thus we get the result

$$
M_{c s}(z)={ }_{i, j}^{\Sigma} \varepsilon_{i} \varepsilon_{j} \quad \mathscr{N}_{i j} W_{i j} \int_{0}^{z} R\left(x, y_{i}\right) R\left(z-x, y_{j}\right) d x
$$

One limitation of this formula is that advance lnowledge of the decay properties (in particular the values of $\mathscr{N}_{\text {ij }}$ ) is required. The spectrum due to random summing events is more difficult to calculate since, due to possible time-separation, the resultant pulse height may be not equal to the sum of individual pulse heights. A method taking this effect into account has been described in detail by Kennett et al ${ }^{(5)}$ and shall not be discussed in this thesis. This random sumaing spectrum is independent of cascades and can be calculated without knowledge of coincidences $\mathscr{N}_{i j}$.

## APPENDIX III

DERIVATION OF INVERSE MATRIX ${ }_{\mathrm{L}}{ }^{-1}$
We wish to find the elements of matrix $I^{-1}$, such that

$$
\underline{L}^{-1} I=I
$$

Since matrix L is upper triangular, the inverse $\mathrm{L}^{-1}$ also has the same property and we can immediately write

$$
e_{i j}^{-1}=0 \text { for } i>j
$$

To find the diagonal elements of $\underline{\underline{L}}^{-1}$ we multiply the $k^{\text {th }}$ row of $\underline{L}^{-I}$ by the $k^{\text {th }}$ column of $I$. The result is

$$
\begin{equation*}
\ell_{k k}^{-1} \ell_{k k}=1 \tag{AS. 3}
\end{equation*}
$$

By definition (2.4.12) for matrix I we can express the elements

$$
\begin{aligned}
\ell_{\mathrm{ldk}} & =\alpha_{k^{2}} \\
\ell_{\operatorname{mn}} & =\frac{\beta_{n}}{n-1} \quad \text { for } n>m_{g} \\
& =0 \quad \text { for } n<m .
\end{aligned}
$$

Consequently, by virtue of (A3.3) we have the diagonal elements

$$
\ell_{k k}^{-1}=\frac{1}{\alpha_{k}} .
$$

It remains to find the upper off-diagonal elements of $\underline{L}^{-1}$. To accomplish this, consider the product of the $k^{\text {th }}$ row of $\underline{I}^{-1}$ by the $(k+j)^{\text {th }}$ column of $I$, where $0<j \leqslant N=k_{0}$ Writing out this product in detail we have

$$
\sum_{i=0}^{j} e_{\mathrm{kk}+i}^{-1} l_{k+i k+j}=0
$$

Substitution of elements (A3.4) in (A3.6) yields the result

$$
{e_{k k+j}^{-1}}_{-1} \frac{-\beta_{k+j}}{\alpha_{k+j}}\left(\sum_{i=j-1)}^{j-1} e_{k k+i}^{-1} .\right.
$$

By using similar arguments we can arrive at the expression

$$
\begin{equation*}
\ell_{k c k+j+1}^{\infty-1}=-\frac{\beta_{k+j+1}}{\alpha_{k+j+1}(k+j)} \sum_{i=1}^{j} \ell_{k k+i}^{-1} \tag{A3. 8}
\end{equation*}
$$

Combining (A3.7) with (A3.8) we can obtain the recursion relation

$$
B_{k j} \equiv \frac{b^{-1} k k+j+1}{b_{k k k+j}^{-1}}=\frac{\beta_{k+j+1}}{\alpha_{k+j+1}(k+j)}\left[\frac{a_{k+j}(k+j-1)}{\beta_{k+j}}-1\right] \cdot A 3.9
$$

Using the quantity $\gamma_{j}$ defined by (2.4.12) the above ratio can be written in the form

$$
\begin{equation*}
R_{k j}=\frac{\left(\gamma_{k+j+1}-1\right)}{\left(r_{k+j}-1\right)} r_{k+j} \tag{A3. 10}
\end{equation*}
$$

We already have determined the diagonal inverse element in the $k^{\text {th }}$ row (see expression (A3.5)). To obtain the next element in this row we multiply the $\mathrm{k}^{\text {th }}$ row of $\mathrm{I}^{-1}$ by the $(\mathrm{k}+1)^{\text {th }}$ column of I . This results in the expression

$$
\begin{equation*}
a_{k k+1}^{-1}=\frac{1}{a_{k}}\left(\gamma_{k+1}-1\right) \tag{A3. 11}
\end{equation*}
$$

By the recursion relation ( $A 3.10$ ) the next inverse element in the $k^{\text {th }}$ row is given by

$$
\begin{equation*}
b_{k k+2}^{-1}=\frac{1}{\alpha_{k}}\left(\gamma_{k+2}-1\right) \gamma_{k+1} . \tag{A3. 12}
\end{equation*}
$$

It is easily seen that by repeated application of the recursion formula we obtain a general inverse matrix element

$$
\begin{equation*}
\ell_{k j}^{-1}=\frac{1}{\alpha_{k}}\left(\gamma_{j}-1\right) \prod_{i=k+1}^{j-1} \gamma_{i} \text { for } k<j \tag{A3. 13}
\end{equation*}
$$

This completes the derivation of expression (2.4.13). ,
It is of interest to derive an effort factor (total number of multiplications and divisions) required for inversion of matrix L. There are ( $N-I$ ) values of $\gamma_{i}$. Each requires one multiplication and one division to evaluate. This results in a total number of operations

$$
\begin{equation*}
M_{\gamma}=2(N-1) \tag{A3. 14}
\end{equation*}
$$

The recursion relation (A3.10) involves a total of (N-2) factors. Since each factor is evaluated by two operations we have a total of operations

$$
\begin{equation*}
M_{R}=2(N-2) \tag{A3. 15}
\end{equation*}
$$

Now we consider the evaluation of inverse elements. Each of the $\frac{N}{2}(N+1)$ non-zero elements requires either one division or one multiplication. Consequently the total number of operations is given by

$$
\begin{align*}
M & =\frac{N^{2}}{2}+\frac{N}{2}+M_{\gamma}+M_{R} \\
& =\frac{N^{2}}{2}+\frac{9 N}{2}-6 .
\end{align*}
$$

This is a considerable improvement over the cubic relation (2.3.14) for the inversion of a general triangular matrix.

When parameters $\alpha$ and $\beta$ are assumed constant the required effort is somewhat less than the effort given by (A3.16). After determination of $\beta / \alpha$ each $\gamma_{i}$ requires only one division. Therefore in this case

$$
M_{\gamma}^{c}=(N-1)+1=N .
$$

Since the $\gamma_{i}$ 's are different the number $M_{R}$ remains the same as in (A3.15).

The diagonal elements are all equal to $1 / \alpha$ and thus require only one division. All offodiagonal elements are obtained by $\mathbb{N}(\mathbb{N}-1) / 2$ multiplications. Thus the total number of operations is

$$
\begin{aligned}
M^{c} & =N+1+\frac{N^{2}}{2}-\frac{N}{2} \\
& =\frac{N^{2}}{2}+\frac{N}{2}+1 .
\end{aligned}
$$

$$
\text { A3. } 18
$$

## APPENDIX IV

FOURIER TRANSFORMS AND CONVOLUTION

First we prove the convolution theorem, which is used in section 2.5. We have the convolution expression (2.5.4a) in the $x$-domain and wish to derive the product (2.5.4b) in the $\omega$-domain. By substituting the Fourier transforms under the convolution integral we can write

$$
\begin{aligned}
F_{3}(x) & =\frac{1}{(2 \pi)^{2}} \iiint_{-\infty}^{\infty} f_{1}\left(\omega_{1}\right) e^{i \omega_{1}(x-y)} f_{2}\left(\omega_{2}\right) e^{i \omega_{2} y} d y d \omega_{1} d \omega_{2} \\
& =\frac{1}{2 \pi} \quad \int_{-\infty}^{\infty} \int_{1}\left(\omega_{1}\right) f_{2}\left(\omega_{2}\right) e^{i \omega_{1} x}\left[\frac{1}{2 \pi} \int_{-\infty}^{\infty} e^{i\left(\omega_{2}-\omega_{1}\right) y} d y\right] d \omega_{1} \cdot d \omega_{2} .
\end{aligned}
$$

The quantity inside the brackets is just one form of the definition (35) for the delta function $\delta\left(\omega_{2}-\omega_{1}\right)$. Thus we can make the substitution $\omega_{2}=\omega_{1}$ and drop the integration over $\omega_{2}$. This leads to the expression

$$
\begin{equation*}
F_{3}(x)=\frac{1}{2 \pi} \quad \int_{-\infty}^{\infty} f_{1}\left(\omega_{1}\right) f_{2}\left(\omega_{1}\right) e^{i \omega_{1} x} d \omega_{1} \tag{A 4.1}
\end{equation*}
$$

Therefore $F_{3}(x)$ is the Fourier transform of $f_{1}(\omega) f_{2}(\omega)$; but by definition it is also the same Fourier transform of $f_{3}(\omega)$. Hence we can write the equality

$$
f_{3}(\omega)=f_{1}(\omega) f_{2}(\omega) .
$$

When we have convolution in the $\omega$-domain (as in (2.5.5a)), it is possible to arrive at the product expression (2.5.5b) by using arguments identical to the ones presented above. .This would complete the proof of both aspects of the convolution theorem.

Next we derive the Fourier transforms of some expressions used in section 2.5. Consider the Gaussian function of (2.5.9a). Its Fourier transform is given by

$$
r(\omega)=\frac{1}{\sqrt{2 \pi} \sigma} \cdot \int_{-\infty}^{\infty} \exp \left[-\frac{x^{2}}{2 \sigma^{2}}-i \omega x\right] d x
$$

We can complete the square in the exponent and write it in the form

$$
-\left(\frac{x}{\sqrt{2} \sigma}+\frac{i \omega \sigma}{\sqrt{2}}\right)^{2}-\frac{\omega^{2} \sigma^{2}}{2}
$$

Then, upon introducing the transformation

$$
y=\frac{x}{\sqrt{2} \sigma}+\frac{i \omega \sigma}{\sqrt{2}}
$$

we can rewrite expression (A4.3),

$$
r(\omega)=\frac{e^{-\frac{\sigma^{2} \omega^{2}}{2}}}{\sqrt{\pi}} \int_{-\infty}^{\infty} e^{-y^{2}} d y \quad \text { A }
$$

The definite integral has the numerical value $\sqrt{\pi}$. Hence the Fourier transform of the Gaussian response function (2.5.9a) is given by

$$
r(\omega)=e^{-\frac{\sigma^{2} \omega^{2}}{2}}
$$

Consider now the set of delta functions

$$
D(x)=\sum_{n=-\infty}^{\infty} \delta(x-n)
$$

Since this set forms a periodic structure symmetrical about the point $\mathbf{x}=0_{8}$ we can expand it in a Fourier cosine series,

$$
D(x)=a_{0}+\sum_{n=1}^{\infty} a_{n} \cos 2 \pi n x
$$

We have

$$
\begin{aligned}
a_{0} & =\int_{-\frac{1}{2}}^{\frac{1}{2}} D(x) d x=1 \\
a_{n} & =2 \int_{-\frac{1}{2}}^{\frac{1}{2}} D(x) \cos 2 \pi n x d x \\
& =2 \int_{-\frac{1}{2}}^{\frac{1}{2}} \delta(x) \cos 2 \pi a x x \\
& =2
\end{aligned}
$$

$$
\text { AU. } 8
$$

Therefore we can write

$$
D(x)=1+2 \sum_{n=1}^{\infty} \cos 2 \pi n x_{0}
$$

Taking the Fourier transform of $D(x)$ we obtain

$$
\begin{aligned}
d(\omega) & =\int_{-\infty}^{\infty} e^{-i \omega x} d x+\sum_{n=1}^{\infty} \int_{-\infty}^{\infty}\left(e^{i 2 \pi n x}+e^{-i 2 \pi n x}\right) e^{-i \omega x} d x \\
& =\sum_{n=\infty}^{\infty} \int_{\infty}^{\infty} e^{i x(2 \pi n-\omega)} d x \\
& =2 \pi \sum_{n=\infty}^{\infty} 6(\omega-2 \pi n)
\end{aligned}
$$

$$
\text { AL. } 10
$$

which establishes the relation (2.5.10b).
Now consider the channel profile $P(x)$ as defined in expression
(2.5.28a). The Fourier transform is

$$
p(\omega)=\int_{-\frac{1}{2}}^{\frac{9}{2}} e^{-i \omega x} d x
$$

According to de Moivre's theorem, we have

$$
e^{-i \omega x}=\cos \omega x-i \sin \omega x \text {. }
$$

Being assymetrical, the second term has no contribution to the integral in ( 84.11 ); bence we can write

$$
\begin{align*}
p(\omega) & =\int_{-\frac{1}{2}}^{\frac{1}{2}} \cos \omega x d x \\
& =\frac{2}{\omega} \sin \left(\frac{\omega}{2}\right) \tag{A4. 12}
\end{align*}
$$

which proves the relation (2.5.28b).

It remains now to derive the result (2.6.3) for "partial inversion" matrix elements. We start by considering again the convolution integral (2.5.1),

$$
\begin{equation*}
M(x)=\quad \int_{-\infty}^{\infty} R(x-y) M(y) d y . \tag{A4. 13}
\end{equation*}
$$

However, this time we try to find a "partial inverse" function $R_{p i n}(x)$ such that the unfolded spectrum retains a resolution function $R_{n}(x)$, which is narrower than the original response $R(x)$. Therefore, instead of (2.5.2), we must write

$$
\int_{-\infty}^{\infty} R_{n}(x \infty y) M(y) d y=\int_{-\infty}^{\infty} R_{p i n}(x-y) M(y) d y
$$

Using the convolution theorem we can express these two equations in the $\omega=$ domain,

$$
\begin{equation*}
m(\omega)=r(\omega) t(\omega) \tag{A4. 15}
\end{equation*}
$$

and

$$
x_{\mathrm{n}}(\omega) t(\omega)=r_{\text {pin }}(\omega) m(\omega)
$$

$$
\text { AL. } 16
$$

These two relations are satisfied when

$$
\begin{equation*}
x_{\text {pin }}(\omega)=\frac{r_{n}(\omega)}{r(\omega)} \tag{AU. 17}
\end{equation*}
$$

When $R(x)$ was digitized by the factor $D(x)$ we had the result

$$
\begin{equation*}
r^{d}(\omega)=\sum_{n=-\infty}^{\infty} r(\omega-2 \pi n) \tag{AU. 18}
\end{equation*}
$$

as given by (2.5.13). In a similar way we digitize now the function $R_{n}(x)$, and obtain

$$
\begin{equation*}
r_{n}^{d}(\omega)=\sum_{n=-\infty}^{\infty} r_{n}(\omega-2 \pi n) \tag{AL. 19}
\end{equation*}
$$

After these digitizations we replace expression (A4.17) by

$$
\begin{equation*}
r_{\text {pin }}(\omega)=\frac{r_{n}^{d}(\omega)}{r^{d}(\omega)} \tag{A 4.20}
\end{equation*}
$$

Following the same procedure as in section 2.5, we define a truncated function

$$
\begin{aligned}
f_{n}^{\prime}(\omega) & =x_{\text {pin }}(\omega) & \text { for }|\omega|<\pi \\
& =\frac{1}{2} r_{\text {pin }}(\omega) & \text { for }|\omega|=\pi \\
& =0, & \text { for }|\omega|>\pi
\end{aligned}
$$

which can be used to obtain the result

$$
\begin{gather*}
R_{\text {pin }}(x)=F_{n}(x) D(x)  \tag{At. 22}\\
F_{n}(x)=\frac{1}{2 \pi} \int_{-\pi}^{\pi} \frac{e^{i \omega x}}{r^{\alpha}(\omega)} r_{n}^{d}(\omega) d \omega
\end{gather*}
$$

where

If we now substitute (A4.22) in expression (A4.14) and use the digitized expression for $R_{n}(x)$, we have

$$
\sum_{m a=-\infty}^{\infty} R_{n}(m) T(x-m)=\sum_{m=-\infty}^{\infty} F_{n}(m) M(x-m)
$$

By using the transformation $x-\infty=k$ this last expression can be rewritten
as

$$
\begin{equation*}
\sum_{k=-\infty}^{\infty} R_{n}(x-k) T(k)=\sum_{k=\infty}^{\infty} F_{n}(x-k) M(k) . \tag{AU. 25}
\end{equation*}
$$

This is analogous to a matrix equation

$$
B_{n} T=F_{n} \underline{M}
$$

$$
\text { AU. } 26
$$

where $R_{n}$ is a matrix containing the reduced width response functions and $\mathbb{F}_{\mathrm{n}}$ is a "partial inverse" matrix with elements

$$
\left(f_{n}\right)_{k l}=\frac{1}{2 \pi} \int_{-\pi}^{\pi} \frac{e^{i \omega(k-l)}}{r^{d}(\omega)} r_{x}^{d}(\omega) d \omega
$$

The relation (2.6.3) is thus established.

## APPENDIX V

## SUPPLEMENTARY PROOFS FOR SECTION 3.3

It is required to prove that the design matrix ${\underset{-}{2}}_{2}$, as defined by (3.3.21), has an inverse $\underline{B}_{2}^{-1}$, provided the response functions $X_{i}(x)$ and $Y_{j}(y)$ form two sets of linearly independent vectors. The proof will be complete if we can show that the vectors formed by the columns of matrix Z are linearly independent. We start by assuming that the columns are linearly dependent and then proceed to show that this leads to an imposesible condition. The linear dependence in Z implies that we can find a non-zero vector

$$
\underline{v}=\left\|\left\{\left\|\left\{v_{i j}\right\}_{j, 1}\right\|^{n_{y}, 1}\right\}_{i, 1}\right\|^{n_{x}, 1}
$$

such that

$$
\underline{\underline{V}}=\left\|\{0\}_{i, 1}\right\|^{N_{X} N_{x}, 1} \ldots \quad A 5.2
$$

It is possible to write the null-vector $\underline{Z} \underline{V}$ in the more detailed form

$$
\begin{align*}
& \|\left\{\begin{array}{l}
n_{x} \\
\sum_{i=1}
\end{array}\left\|\left\{x_{i}(x) y_{j}(y)\right\}_{y, j}\right\|^{N_{y}, n_{y}}\left\|\left\{_{\left.v_{i j}\right\}_{j, 1}} \|^{n_{y}, 1}\right\} x, 1\right\|^{N_{x}, I}\right. \\
& =0 \tag{AS. 3}
\end{align*}
$$

If we look only at the subvectors indexed by $x$ we can say that

$$
\left\|\left\{y_{j}(y)\right\}_{y, j}\right\|^{N_{y}, n_{y}}\left\|\left\{\sum_{i=1}^{n_{x}} x_{i}(x) v_{i j}\right\}_{j, I}\right\|^{n_{y}, 1}=0^{N_{y}, 1}
$$

for every $x=1,2, \ldots, N_{x}$. The second factor in equation (A5.4) cannot be a null vector for every $x$ since $\underline{V}$ is not a null vector and the $X_{i}(x)$ are linearly independent. But then equation (A5.4) implies that the $Y_{j}(y)$ are linearly dependent, which is not true. Hence we must reject the assumption that the columns of $\underset{Z}{ }$ are linearly dependent. According to Cramer (49) the linear independence in $\underset{Z}{Z}$ is sufficient to insure that the inverse matrix $\mathrm{B}_{2}^{-1}$ exists.

Next we turn to the derivation of expression (3.3.39) which gives the full covariance matrix of estimate vector $A_{2}^{\prime \prime}$. We join all data vectors (3.3.25) to form one long vertical vector defined by

$$
\underline{M}_{0}^{\prime}=\left\|\left\{\left\|\left\{M_{2}^{\prime}(x, y)\right\}_{x, 1}\right\|^{x^{\prime}, 1}\right\}_{y, 1}\right\|^{N_{y}, I}
$$

Likewise, we string out the vectors of intermediate coefficients by defining

$$
Q^{\prime}=\left\|\left\{\left\|\left\{q_{i}^{\prime}(y),\right\}_{i, 1}\right\|^{n_{x}, 1}\right\}_{y, 1}\right\|^{N_{y}, 1}
$$

Making use of these definitions we propose to find an $n_{x} n_{y}$ by $N_{x} N_{y}$ size matrix $\underline{P}$ such that

$$
\underline{A}_{2}^{\prime \prime}=\underline{P} \underline{M}_{0}^{\prime} \cdot
$$

According to a theorem ${ }^{(50)}$ mentioned in the text the required covariance matrix is then expressible in the form

$$
\underline{\underline{G}}_{A_{2}^{\prime \prime}}=\underline{\mathrm{P}}_{\mathrm{C}_{\mathrm{M}}^{1}} \underline{\mathrm{P}}^{\mathrm{T}}
$$

Equation (3.3.29) gives $N_{y}$ least squares solutions in the $x$ dimension. These solutions can be grouped together and expressed in one matrix equation

$$
\begin{equation*}
\underline{Q}^{\prime}=\underline{H}_{0} \underline{M}_{0}^{\prime} \text {, } \tag{AS. 9}
\end{equation*}
$$

where

$$
\underline{H}_{0}=\left\|\left\{\left\|\underline{F}_{-1}^{-1} \underline{x}^{T} \underline{W}_{y}\right\|^{n_{x}, N_{x}} \quad \begin{array}{lll} 
& \delta_{y i}
\end{array}\right\}_{y, i}\right\|^{N_{y}, N_{y}} . \quad A 5.10
$$

Before we can group solutions (3.3.35) for the $y$-dimension, we must rearrange the intermediate vector $Q^{\prime}$ so that it consists of a string of vectors $Q_{i}^{\prime}$. This is accomplished by a rearranging matrix $R_{o}$ such that

$$
\begin{align*}
Q_{0}^{\prime} & =\left\|\left\{\left\|\left\{a_{i}^{\prime}(y)\right\}_{y, 1}\right\|^{N_{y}, 1}\right\}_{i, 1}\right\|^{n_{x}, 1} \\
& =R_{0} Q^{\prime} .
\end{align*}
$$

The desired rearranging matrix can be defined by

$$
R_{o}=\left\|\left\{\left\|\left\{\delta_{i j} \delta_{k y}\right\}_{k, i}\right\|^{N_{y}, n_{x}}\right\}_{j, y}\right\|^{n_{x}, N_{y}}, \quad A 5.12
$$

as substitution in (A5.11) will readily show.
After introducing one more definition,

$$
\underline{R}_{1}=\left\|\left\{\| \|_{i}^{-1} \underline{\underline{x}}^{T} \underline{W}_{i} \|^{n^{n}, N_{y}} \quad \begin{array}{ll} 
& \delta_{i j}
\end{array}\right\}_{i, j}\right\|^{n_{x}, n_{x}}, \quad A 5.13
$$

we can write down the final solution vector

$$
\begin{equation*}
\underline{A}_{2}^{\prime \prime}=\underline{R}_{1} R_{0} H_{0} M_{0}^{\prime} . \tag{AF. 14}
\end{equation*}
$$

Thus the covariance matrix of estimate $A_{2}^{\prime \prime}$ is given by
which reduces to expression (3.3.39) after a certain amount of matrix algebra.

It remains now to prove the inequalities in expression (3.3.45). It is shown by Cramér (49) that the reciprocal of a diagonal element in the positive definite design matrix is not greater than the value of corresponding diagonal element in the inverse matrix. This fact immediately establishes the inequality

$$
\begin{equation*}
D\left(a_{k \ell}^{\prime}\right) \geqslant{\frac{I}{\left\{\underline{B}_{2}\right\}}}_{i, i}, \tag{A5. 16}
\end{equation*}
$$

where $i=(k-1) n_{y}+h_{0}$
Since the diagonal submatrix $B_{-k k}$ is also positive definite, we have the second inequality

$$
\left\{\underline{B}_{k k k}^{-1}\right\}_{\ell, l} \geqslant \frac{I}{\left\{\underline{B}_{2}\right\}_{i_{0} i}} .
$$

Next we consider the inequality between expressions $D\left(a_{K, l}^{0}\right)$ and $\left\{\underline{B}_{\mathrm{Kck}}^{-1}\right\}_{\mathrm{LL}}$. Matrix ${\underset{-2}{ }}$ can be partitioned according to the following scheme:

In a similar way we partition the inverse matrix ${\underset{-}{-1}}_{-1}^{2}$ and label the submatrices by $\mathrm{C}_{i j}{ }^{\text {• }}$ Upon multiplication of these partitioned matrices we can write

$$
\begin{align*}
& \underline{A}_{11} \underline{C}_{11}+\underline{A}_{12} \underline{C}_{21}=\underline{I}^{n_{y}}, \\
& \underline{C}_{21} \underline{A}_{11}+\underline{C}_{22} \underline{A}_{21}=\underline{0},  \tag{AS. 20}\\
& \underline{C}_{21} \underline{A}_{12}+\underline{C}_{22} \underline{A}_{22}=\underline{I}^{\left(n^{-1) n_{y}}\right.}, \tag{AS. 21}
\end{align*}
$$

where $I$ is a diagonal identity matrix (with rank indicated by the superscript). From (A5.19) we have

$$
\begin{equation*}
\underline{C}_{11}=A_{11}^{-1}-\underline{A}_{11}^{-1} A_{12} \underline{C}_{21}, \tag{AF. 22}
\end{equation*}
$$

and from (A5.20) we obtain

$$
\begin{equation*}
\underline{C}_{21}=-\underline{C}_{22} \mathbb{A}_{21} A_{11}^{-1} \tag{AF. 23}
\end{equation*}
$$

These last two expressions can be combined to yield the result

$$
\begin{equation*}
\underline{G}_{11}=A_{11}^{-1}+A_{11}^{-1} A_{12} \underline{G}_{22} A_{21} A_{11}^{-1} \cdot \tag{AF. 24}
\end{equation*}
$$

We now let the positive definite matrix $\underline{G}_{22}$ have elements $c_{i j}$ and define another set of elements $t_{i j}$ such that

$$
\begin{equation*}
A_{11}^{-1} A_{12}=\left\|\left\{t_{i j}\right\}_{i, j}\right\|^{n_{y},\left(n_{x}-1\right) n_{y}} \tag{AS. 25}
\end{equation*}
$$

The second term in (A5.24) can now be written in the form

$$
\begin{gathered}
\left\|\left\{t_{i j}\right\}_{i, j}\right\|\left\|^{n_{y}\left(n_{x}-1\right) n_{y}}\right\|\left\{c_{i j}\right\}_{i, j}\left\|^{\left(n_{x}-1\right)_{y}\left(n_{x}-1\right) n_{y}}\right\|\left\{t_{i j}\right\}_{j, I} \|^{\left(n_{x}-1\right) n_{y}, n_{y}} \\
=\left\|\left\{\begin{array}{cc}
\begin{array}{c}
\left.n_{x}-1\right) n_{y} \\
\sum_{k, j=1}
\end{array} t_{\ell k} & c_{k j} \\
t_{i j}
\end{array}\right\}_{\ell, i}\right\|^{n_{y}, n_{y}} \quad .
\end{gathered}
$$

The $\ell^{\text {th }}$ diagonal element of this matrix is given by

$$
\sum_{k, j=1}^{\left(n_{x}-1\right) n_{y}} \quad t_{\ell k} c_{k j}{ }^{t}{ }_{l j},
$$

which is a positive definite quadratic form by virtue of matrix ${\underset{-}{22}}$ being
positive definite. Therefore, going back to equation (A5.24), we see that the diagonal elements of $\underline{C}_{11}$ are greater than or equal to the corresponding diagonal elements in $A_{11}^{-1}$. This establishes the inequality

$$
D\left(a_{k \ell l}^{\prime}\right) \geqslant\left\{{\underset{-}{k k}}_{-1}^{{ }_{k k}}\right\}_{\ell, \ell}
$$

for the index value $k=1$. Similar arguments can be used to prove the inequality for other values of $k$.

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